

Rigorous Fusion of Gravity Field, Altimetry and Stationary Ocean Models

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Abstract

Many characteristics of the ocean circulation are reflected in the mean dynamic topography (MDT). Therefore observing the MDT provides valuable information for evaluating or improving ocean models. Using this information is complicated by the inconsistent representation of MDT in observations and ocean models. This problem is addressed by a consistent treatment of satellite altimetry and geoid height information on an ocean model grid. The altimetric sea surface is expressed as a sum of geoid heights represented by spherical harmonic functions and the mean dynamic topography parameterized by a finite element method. Within this framework the inversion and smoothing processes are avoided that are necessary in step-by-step approaches, such that the normal equations of the MDT can be accumulated in a straightforward way. Conveniently, these normal equations are the appropriate weight matrices for model-data misfits in least-squares ocean model inversions.

Two prototypes of these rigorously combined MDT models, with an associated complete error description including the omission error, are developed for the North Atlantic Ocean and assimilated into a 3D-inverse ocean model.

The ocean model solutions provide evidence that satellite observations and oceanographic data are consistent within prior errors.

Keywords: mean dynamic topography, gravity field, altimetry, inverse ocean model, combined model

1. Introduction

The ocean's mean dynamic topography (MDT) contains valuable information about the ocean circulation (Wunsch and Stammer, 1998). Therefore, estimates of the MDT have the great potential of improving ocean circulation estimates when properly combined with other information, for example, adequate ocean models.

In principle, dynamic topography is the difference between the altimetric mean sea surface and geoid height, but calculating this difference is not straightforward because the data types have different representations and spatial resolutions. Gravity field models derived from satellite missions (e.g. GRACE and GOCE) are usually represented by spherical harmonic functions. The downward continuation process from satellite altitude to the Earth's surface leads to an amplification of high frequencies while small scale signals in the smooth field at satellite altitude are hidden in the measurement noise. As a consequence of this unfavorable signal-to-noise ratio, high degree spherical harmonic coefficients in the geoid computation cannot be separated very well. Therefore, the geoid models are typically truncated in a regularization process at a maximum degree L to yield a band-limited representation. The propagated errors of the truncated model, however, only represent the modeled part of the signals (commission error). The part of

21 the signal for degrees greater than L , that often is omitted, also ought to be
22 taken into account as omission errors to form a consistent model (Losch et al.,
23 2002). In contrast to the geoid information, the altimetric measurements are
24 given as point values or mean values over the footprint of the radar signal
25 along the ground tracks of the satellite. The sample rate along these tracks
26 is very high and the altimetric measurements contain information with high
27 spatial resolution. The sampling is much coarser in the cross track direction,
28 because the ground tracks of the repeating orbit only form a coarse grid.

29 The two data sets cannot be combined in a straightforward way, because
30 their resolution is different in both space and spectral domain. Special filter
31 processes are introduced (Jekeli, 1981, 1996; Wahr et al., 1998; Swenson and
32 Wahr, 2006; Kusche, 2007) to homogenize all the available information with
33 respect to a least common subspace. Only in this subspace, different phe-
34 nomena can be compared and hypotheses can be formulated, but the amount
35 of signal lost in such procedures remains unclear. All derived statements are
36 only valid with respect to this subspace, and the geometrical interpretation
37 of the exact content of these filtered quantities is not straightforward. The
38 inherent restriction of altimetry profiles to the ocean with associated bound-
39 ary issues and the transition from the open ocean to shallow shelves are
40 additional problems that have been identified. For example, Albertella and
41 Rummel (2009) conclude that extending the altimetric data set to the entire
42 globe will inevitably result in a distortion of its spectral content.

43 To overcome this drawback of filter or smoothing processes a rigorous fu-
44 sion of the gravity field, altimetric observations, and stationary ocean models
45 is proposed. In this approach, the altimetric sea surface is interpreted as the

46 sum of geoid heights represented by spherical harmonic functions and the
47 mean dynamic topography (MDT) parameterized by finite elements. With
48 this combined model the normal equations of the MDT are built directly by
49 a Schur decomposition. Subsequently, these normal equations can be used in
50 stationary ocean models to weight the model data misfit in a least-squares
51 sense. Thus, the inversion of potentially rank deficient covariance matrices
52 and additional smoothing processes that are necessary in step-by-step
53 approaches are avoided.

54 The paper is organized as follows. In section 2 the interface for the assim-
55 ilation of the MDT into ocean circulation models is defined and the deter-
56 ministic and stochastic MDT models are discussed. In section 3 the data sets
57 used in the numerical studies are introduced and first results of the behavior
58 of the estimated MDTs and their accuracy are given. Section 4 shows the
59 results of the integration of the MDT normal equations in the ocean circu-
60 lation model and discusses the effects of the new data sets on characteristics
61 of the ocean model such as temperature distribution, meridional overturning
62 and heat transports. Conclusions are drawn in section 5.

63 **2. Methodology**

64 In theory, the MDT is the mean sea surface (MSS) referenced to the geoid
65 (N) but a simple combination of MSS and N data is not straightforward due
66 to the different spatial resolutions and representations of altimetric measure-
67 ments and a geoid model. Here, the altimetric mean sea surface is interpreted
68 as the sum of geoid height and mean dynamic topography

$$69 \quad \text{MSS}(\phi, \lambda) = \text{N}(\phi, \lambda) + \text{MDT}(\phi, \lambda) \quad (1)$$

70 where ϕ , and λ are the spherical geocentric coordinates. The geoid is repre-
 71 sented as a sum of spherical harmonic functions

72

73 $N(\phi, \lambda) =$

74
$$\frac{GM}{R\gamma(B)} \sum_{n=0}^{\infty} \sum_{m=0}^n \left(\frac{R}{r}\right)^{n+1} \bar{P}_{nm}(\cos \phi) (\bar{C}_{nm} \cos(m\lambda) + \bar{S}_{nm} \sin(m\lambda)) \quad (2)$$

75

76 with the radius vector r depending only on latitude ϕ , the Earth's radius R ,
 77 the gravitational constant times the Earth's mass GM , the geodetic latitude
 78 B , the normal gravity γ , the fully normalized Legendre functions $\bar{P}_{nm}(\cos \phi)$
 79 and the Stokes coefficients \bar{C}_{nm} and \bar{S}_{nm} .

80 The mean dynamic topography is represented by a linear combination of
 81 finite element basis functions $b_k(\phi, \lambda)$, $k \in K$ with a set of indices K that
 82 label the basis functions

83
$$\text{MDT}(\phi, \lambda) = \sum_{k \in K} a_k b_k(\phi, \lambda) . \quad (3)$$

84 In this study we use linear polynomials as basis functions b_k . The definition of
 85 the finite elements, that is, the choice of the basis functions and nodal points,
 86 corresponds directly to the ocean circulation model used in section 4. In this
 87 way the mean dynamic topography can be assessed directly on the target
 88 grid; the coefficients a_k represent the nodal values of the field of interest.

89 Assembling the spherical harmonic coefficients \bar{C}_{nm} and \bar{S}_{nm} in the vector
 90 of unknowns \mathbf{x}_{cs} and the coefficients of the linear combination of the finite
 91 elements in \mathbf{x}_{FE} , the observation equations for the altimetric information
 92 \mathbf{l}_{MSS} become

93
$$\mathbf{l}_{\text{MSS}} + \mathbf{v}_{\text{MSS}} = \begin{bmatrix} \mathbf{A}_{cs} & \mathbf{A}_{FE} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs} \\ \mathbf{x}_{FE} \end{bmatrix} \quad (4)$$

94 with the error covariance matrix Σ_{MSS} . Then the normal equations for the
 95 mean sea surface are constructed as

$$96 \begin{bmatrix} \mathbf{A}_{cs}^T \Sigma_{\text{MSS}}^{-1} \mathbf{A}_{cs} & \mathbf{A}_{cs}^T \Sigma_{\text{MSS}}^{-1} \mathbf{A}_{FE} \\ \mathbf{A}_{FE}^T \Sigma_{\text{MSS}}^{-1} \mathbf{A}_{cs} & \mathbf{A}_{FE}^T \Sigma_{\text{MSS}}^{-1} \mathbf{A}_{FE} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs} \\ \mathbf{x}_{FE} \end{bmatrix} = \begin{bmatrix} \mathbf{A}_{cs}^T \Sigma_{\text{MSS}}^{-1} \mathbf{l}_{\text{MSS}} \\ \mathbf{A}_{FE}^T \Sigma_{\text{MSS}}^{-1} \mathbf{l}_{\text{MSS}} \end{bmatrix}, \quad (5)$$

98 and in abbreviated form

$$99 \begin{bmatrix} \mathbf{N}_{cs}^{\text{MSS}} & \mathbf{N}_{cs,FE}^{\text{MSS}} \\ \mathbf{N}_{FE,cs}^{\text{MSS}} & \mathbf{N}_{FE}^{\text{MSS}} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs} \\ \mathbf{x}_{FE} \end{bmatrix} = \begin{bmatrix} \mathbf{n}_{cs}^{\text{MSS}} \\ \mathbf{n}_{FE}^{\text{MSS}} \end{bmatrix}. \quad (6)$$

101 In this study a static solution of a satellite-derived gravity field model from
 102 GRACE or GOCE is used, for which the Stokes coefficients \bar{C}_{nm} , \bar{S}_{nm} and the
 103 full variance/covariance matrix Σ_{cs}^{G} are available, and therefore the normal
 104 equations

$$105 \mathbf{N}_{cs}^{\text{G}} \mathbf{x}_{cs} = \mathbf{n}_{cs}^{\text{G}}. \quad (7)$$

106 In general, the normal equations of altimetric measurements (6) and the
 107 normal equations of the geoid model (7) result from independent observation
 108 groups, so that the summation theorem of normal equations can be applied
 109 to give

$$110 \begin{bmatrix} \mathbf{N}_{cs}^{\text{G}} + \mathbf{N}_{cs}^{\text{MSS}} & \mathbf{N}_{cs,FE}^{\text{MSS}} \\ \mathbf{N}_{FE,cs}^{\text{MSS}} & \mathbf{N}_{FE}^{\text{MSS}} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs} \\ \mathbf{x}_{FE} \end{bmatrix} = \begin{bmatrix} \mathbf{n}_{cs}^{\text{G}} + \mathbf{n}_{cs}^{\text{MSS}} \\ \mathbf{n}_{FE}^{\text{MSS}} \end{bmatrix}. \quad (8)$$

111 Eliminating the gravity field parameters \mathbf{x}_{cs} from these normal equations by
 112 using a Schur decomposition (Golub and van Loan, 1983, page 192) provides
 113 the normal equations for the mean dynamic topography, in short

$$114 \bar{\mathbf{N}}_{FE} \mathbf{x}_{FE} = \bar{\mathbf{n}}_{FE}. \quad (9)$$

115 These normal equations form a consistent set of information. If all pa-
 116 rameters are determined by the observations it is straightforward to solve the

117 system and compute the covariance matrix. If only some of the parameters
 118 are determined by the observations the system becomes unstable or singular.
 119 But this is harmless for the approach in this paper, because an inversion of
 120 the normal equation matrix is not required. Instead the normal equations
 121 are used directly as weights for the new MDT in the Inverse Finite Element
 122 Ocean circulation Model (IFEOM). IFEOM is a stationary inverse model
 123 that solves the minimization problem

$$124 \quad J = \frac{1}{2} \sum_i J_i \stackrel{!}{=} \min \quad (10)$$

125 that is subject to stationary balances of ocean momentum, energy (potential
 126 temperature), salt and mass. The cost function (10) contains contributions
 127 from quadratic model-data differences (temperature and salinity from a hy-
 128 drographic atlas and MDT) weighted by the inverses of their respective er-
 129 ror covariances. The contributions J_i can also be prior information such as
 130 smoothness of the solution. For the relative weighting of the different cost
 131 function terms, the hydrographic data is scaled by their annual variance.
 132 The resulting weights typically increase with depth where the ocean tends
 133 to be quiescent. Towards the open boundary at 4.5°N, weights are increased
 134 in order to constrain the model solution to the first guess in the absence of
 135 better information. As the gain of information by the new MDT and its error
 136 covariance matrix is to be assessed, all these weights remain unchanged in
 137 our experiments to allow for comparison. Details of IFEOM can be found in
 138 Sidorenko (2004) and Richter (2010).

139 In general, the error correlations of the observations are unknown a pri-
 140 ori so that most covariances reduce to diagonal matrices. Here IFEOM is

141 extended by taking into account the full MDT error covariances

$$\begin{aligned}
 142 \quad J_{MDT} &= (\mathbf{x}_{FE}^{\text{data}} - \mathbf{x}_{FE}^{\text{model}})^T \boldsymbol{\Sigma}_{\mathbf{x}_{FE}}^{-1} (\mathbf{x}_{FE}^{\text{data}} - \mathbf{x}_{FE}^{\text{model}}) \\
 &= (\mathbf{x}_{FE}^{\text{data}} - \mathbf{x}_{FE}^{\text{model}})^T \bar{\mathbf{N}}_{FE} (\mathbf{x}_{FE}^{\text{data}} - \mathbf{x}_{FE}^{\text{model}})
 \end{aligned} \tag{11}$$

143 with $\mathbf{x}_{FE}^{\text{data}}$ being the ‘‘observed’’ data derived from gravimetry and altime-
 144 try and $\mathbf{x}_{FE}^{\text{model}}$ being their modeled counterparts. The estimation procedure
 145 requires the inverse of the variance/covariance matrix $\boldsymbol{\Sigma}_{\mathbf{x}_{FE}}$ for weighting
 146 the model-data misfit. This inverse is exactly equal to the normal equation
 147 matrix $\bar{\mathbf{N}}_{FE}$, so that the normal equations are directly used within IFEOM.

148 Unfortunately, the data sets in question are not homogeneous: The alti-
 149 metric mean sea surface has a spatial resolution that is much higher than that
 150 of the geoid model; the spatial resolution of the geoid is homogeneous over
 151 the globe but the altimetric measurements are only available on the tracks over
 152 the ocean. Therefore the frequency spectrum is split into different domains
 153 by the individual observations. This separation is described in the following
 154 section and special attention is paid to the infinite-dimensional parameter
 155 space of the Stokes coefficients.

156 *2.1. Observation equations*

157 In this study the static solution of the latest University-of-Bonn GRACE-
 158 only gravity field model ITG-Grace2010s (Mayer-Gürr et al., 2010) is used. It
 159 is available up to degree and order 180, corresponding to a half-wavelength of
 160 111 km, with the full variance/covariance information $\boldsymbol{\Sigma}_{ITG}$. Geoid heights
 161 are as accurate as 1 cm at degree and order 150. Consequently, $L = 150$
 162 is chosen to divide the vector of unknowns $\mathbf{x}_{cs}^{\text{GRACE}}$ into \mathbf{x}_{cs1} representing
 163 the spherical harmonics up to $L = 150$ and the remaining less accurate

164 parameters \mathbf{x}_{cs_2} for degrees 151 to 180 so that

$$165 \quad \begin{bmatrix} \mathbf{x}_1^{\text{ITG}} \\ \mathbf{x}_2^{\text{ITG}} \end{bmatrix} + \begin{bmatrix} \mathbf{v}_1^{\text{GRACE}} \\ \mathbf{v}_2^{\text{GRACE}} \end{bmatrix} = \begin{bmatrix} \mathbf{I} & \mathbf{0} \\ \mathbf{0} & \mathbf{I} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs_1} \\ \mathbf{x}_{cs_2} \end{bmatrix}. \quad (12)$$

166 GRACE measurements are assumed not to contribute to the signal beyond
 167 degree and order 180 in this study. The mean sea surface is modeled by four
 168 groups of spherical harmonics and one for the finite elements. The param-
 169 eter groups \mathbf{x}_{cs_1} , \mathbf{x}_{cs_2} together with the finite elements determine the lower
 170 frequencies in analogy to the geoid coefficients. The parameter group \mathbf{x}_{cs_3}
 171 describes a transition domain between the observed and the truncated (omit-
 172 ted) spherical harmonic spectrum. The infinite group \mathbf{x}_{cs_4} is determined by
 173 additional external information (cf. section 2.2). Frequencies $>$ degree 180
 174 are only taken into account in the parameterization of the mean sea surface.
 175 In this study the mean sea surface is expanded as a sum of spherical harmon-
 176 ics up to degree and order 240. This choice determines the third group of
 177 spherical harmonics \mathbf{x}_{cs_3} . The particular choice of this domain is somewhat
 178 arbitrary. Tuning showed that the choice of 240 gave reasonable results and
 179 that these results are robust to small variations of this maximum degree.
 180 The infinite set of coefficients beyond 240 as a fourth set \mathbf{x}_{cs_4} completes the
 181 parameter vector. The full representation of MSS is

$$182 \quad \mathbf{l}_{\text{MSS}} + \mathbf{v}_{\text{MSS}} = \begin{bmatrix} \mathbf{A}_{cs_1} & \mathbf{A}_{cs_2} & \mathbf{A}_{cs_3} & \mathbf{A}_{cs_4} & \mathbf{A}_{FE} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs_1} \\ \mathbf{x}_{cs_2} \\ \mathbf{x}_{cs_3} \\ \mathbf{x}_{cs_4} \\ \mathbf{x}_{FE} \end{bmatrix}. \quad (13)$$

183 Gravity field and altimetric observations determine only part of the fre-
 184 quency spectrum, so that additional information based on the smoothness
 185 of the potential (Schuh and Becker, 2010) is introduced. Considering the
 186 normally distributed random variables for the Stokes coefficients $\boldsymbol{x}_{cs}^{\text{smooth}} \sim$
 187 $\mathcal{N}(\mathbf{0}, \boldsymbol{\Sigma}_{cs}^{\text{smooth}})$ results in the following pseudo-observation equations

$$188 \quad \begin{bmatrix} \mathbf{0} \\ \mathbf{0} \\ \mathbf{0} \end{bmatrix} + \begin{bmatrix} \boldsymbol{v}_1^{\text{smooth}} \\ \boldsymbol{v}_2^{\text{smooth}} \\ \boldsymbol{v}_3^{\text{smooth}} \end{bmatrix} = \begin{bmatrix} \boldsymbol{I} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \boldsymbol{I} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \boldsymbol{I} \end{bmatrix} \begin{bmatrix} \boldsymbol{x}_{cs2} \\ \boldsymbol{x}_{cs3} \\ \boldsymbol{x}_{cs4} \end{bmatrix}. \quad (14)$$

189 Kaula's rule (Kaula, 1966) describes the degree-wise signal content of the
 190 gravitational potential coefficients in terms of degree variances

$$191 \quad \sigma_n^2 = 10^{-10} \frac{2n+1}{n^4} \Rightarrow \sigma_{nm}^2 = \frac{10^{-10}}{n^4}. \quad (15)$$

192 Thus, the stochastic model results in

$$193 \quad \boldsymbol{\Sigma}_{cs}^{\text{smooth}} = \begin{bmatrix} \boldsymbol{\Sigma}_{cs2}^{\text{smooth}} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \boldsymbol{\Sigma}_{cs3}^{\text{smooth}} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \boldsymbol{\Sigma}_{cs4}^{\text{smooth}} \end{bmatrix}$$

$$194 \quad = \begin{bmatrix} \text{diag}(\boldsymbol{\sigma}_{nm2}^2) & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \text{diag}(\boldsymbol{\sigma}_{nm3}^2) & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \text{diag}(\boldsymbol{\sigma}_{nm4}^2) \end{bmatrix}. \quad (16)$$

195 Figure 1 summarizes schematically the frequency domains and the associ-
 196 ated parameterizations and accuracies of the individual observation groups.
 197 Note, that the domain of the parameter group \boldsymbol{x}_{cs3} is mainly determined by
 198 the altimetric observations. This domain is called transfer domain and serves
 199 as a buffer between high and low frequency parts of the MSS spectrum. Its

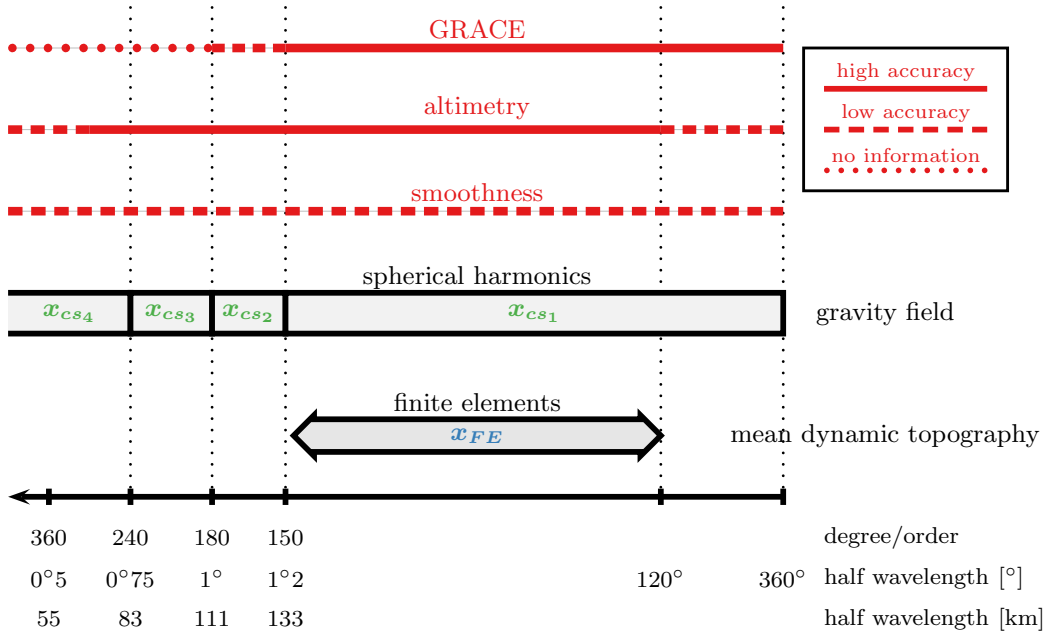


Figure 1: Sketch of frequency domains of different observations their parameterization and accuracy.

200 purpose is to reduce leakage of the high frequencies of the MSS into the com-
 201 mission domain. The high frequency part of the gravity field is not targeted
 202 in this approach.

203 *2.2. Parameterization of the infinite-dimensional space*

204 The altimetric measurements contain frequencies beyond degree and or-
 205 der 240 that ought to be used. The spherical harmonics for this remaining
 206 frequency domain up to infinity are collected in the parameter vector \mathbf{x}_{CS4}

207 and treated separately. Hence, (13) can be recast as

$$208 \quad \mathbf{l}_{\text{MSS}} + \mathbf{v}_{\text{MSS}} = \begin{bmatrix} \mathbf{A}_{cs1} & \mathbf{A}_{cs2} & \mathbf{A}_{cs3} & \mathbf{A}_{FE} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs1} \\ \mathbf{x}_{cs2} \\ \mathbf{x}_{cs3} \\ \mathbf{x}_{FE} \end{bmatrix} + \mathbf{A}_{cs4} \mathbf{x}_{cs4}. \quad (17)$$

209 The extra part of the observation equations $\mathbf{A}_{cs4} \mathbf{x}_{cs4}$ can now be expressed
 210 in terms of random variables

$$211 \quad \mathbf{S} = \mathbf{A}_{cs4} \mathbf{X}_{cs4}. \quad (18)$$

212 The random variable \mathbf{S} is defined by its first two moments, the expectation
 213 $\mathbf{E}\{\mathbf{S}\}$ and covariances $\mathbf{\Sigma}\{\mathbf{S}\}$. In the following, three different choices of the
 214 stochastic characteristics of \mathbf{S} are discussed. They represent two extreme
 215 cases, one in which no or only very little prior information is assumed (Rifu-
 216ugio01 and Rifugio02) and one in which the best possible prior information
 217 about the omitted signal is used. In the latter case the EGM08 (Pavlis et al.,
 218 2008) serves as a place holder or proxy for such information. A realistic
 219 assumption is that the omission error probably lies between these extreme
 220 cases.

221 *2.2.1. Approach 1 – Rifugio01*

222 The first model Rifugio01 assumes no prior information about the sig-
 223 nal content of the gravity field beyond degree and order 240 and empirical
 224 methods are applied to fill the gap. First, the mean sea surface is deter-
 225ministically approximated. After subtracting this trend function from the
 226 mean sea surface the residual signal is analyzed. This results in an empirical

227 auto-covariance function $Cov_{\text{emp}}(\phi, \lambda, \phi', \lambda')$ so that the covariance matrix
 228 $\Sigma_{\mathcal{S}}^{\text{emp}}$ can be assembled. The expectation value of the signal is assumed to
 229 be zero due to the reduction by the deterministic model. The stochastic
 230 characteristics of \mathcal{S} are

$$231 \quad \mathbf{E}\{\mathcal{S}\} = \mathbf{0} := \Delta \mathbf{l}_{\text{MSS}}, \quad \Sigma\{\mathcal{S}\} = \Sigma_{\mathcal{S}}^{\text{emp}} := \Sigma_{\Delta \text{MSS}}. \quad (19)$$

232 2.2.2. Approach 2 – Rifugio02

233 For the model Rifugio02, smoothness of the gravity field according to
 234 Kaula’s rule of thumb is introduced as prior information: $\mathcal{X}_{cs_4} \sim \mathcal{N}(\mathbf{0}, \Sigma_{cs_4}^{\text{smooth}})$.
 235 Because the coefficients $\sigma_{nm}^2 = \frac{1}{2n+1}\sigma_n^2$ are not correlated in the model (see
 236 section 2.1) the covariance in terms of geoid heights can be written as

$$237 \quad Cov(N(\phi, \lambda), N(\phi', \lambda')) =$$

$$238 \quad \frac{G^2 M^2}{R^2 \gamma(B) \gamma(B')} \sum_{n=241}^{\infty} \frac{R^{2(n+1)}}{(rr')^{n+1}} \sigma_n^2 P_n(\cos \psi) \quad (20)$$

239 with the Legendre polynomials $P_n(\cos \psi)$ and the spherical distance ψ . De-
 240 noting the resulting covariance matrix as $\Sigma_{\mathcal{S}}^{\text{Kaula}}$ yields for the stochastic
 241 characteristics

$$242 \quad \mathbf{E}\{\mathcal{S}\} = \mathbf{0} := \Delta \mathbf{l}_{\text{MSS}}, \quad \Sigma\{\mathcal{S}\} = \Sigma_{\mathcal{S}}^{\text{Kaula}} := \Sigma_{\Delta \text{MSS}}. \quad (21)$$

243 2.2.3. Approach 3 – Rifugio03

244 The gravity field model EGM08 (Pavlis et al., 2008) is available to spheri-
 245 cal harmonic degree and order 2160. Here, this information is used to reduce
 246 the mean sea surface by the geoid signal of the EGM08 in the range between
 247 degree and order 241 and 2160. Beyond 2160 the frequencies are treated in
 248 analogy to section 2.2.2. The EGM08 provides error estimates which are used
 249
 250

251 to estimate the accuracy of the geoid height information. Based on the error
 252 degree variances $\sigma_{n,\text{EGM08}}^2$ the overall covariance information can be written
 253 as
 254

$$\begin{aligned}
 255 \quad \text{Cov}(\mathbf{N}(\phi, \lambda), \mathbf{N}(\phi', \lambda')) = \\
 256 \quad \frac{G^2 M^2}{R^2 \gamma(B) \gamma(B')} \sum_{n=241}^{2160} \frac{R^{2(n+1)}}{(rr')^{n+1}} \sigma_{n,\text{EGM08}}^2 P_n(\cos \psi) \\
 257 \quad + \frac{G^2 M^2}{R^2 \gamma(B) \gamma(B')} \sum_{n=2161}^{\infty} \frac{R^{2(n+1)}}{(rr')^{n+1}} \sigma_n^2 P_n(\cos \psi) . \quad (22) \\
 258
 \end{aligned}$$

259 Assembling the geoid height information of the EGM08 in the vector $\Delta \mathbf{l}_{\text{EGM08}}$
 260 and the covariances for the respective frequency domains in the matrices
 261 $\Sigma_{S1}^{\text{EGM08}}$ and $\Sigma_{S2}^{\text{Kaula}}$ leads to

$$\begin{aligned}
 262 \quad \mathbf{E}\{\mathcal{S}\} &= \Delta \mathbf{l}_{\text{EGM08}} := \Delta \mathbf{l}_{\text{MSS}}, \\
 263 \quad \Sigma\{\mathcal{S}\} &= \Sigma_{S1}^{\text{EGM08}} + \Sigma_{S2}^{\text{Kaula}} := \Sigma_{\Delta \text{MSS}} . \quad (23)
 \end{aligned}$$

264 2.3. The model

265 Finally, the complete observation equations for the altimetric measure-
 266 ments are

$$267 \quad \bar{\mathbf{l}}_{\text{MSS}} + \mathbf{v}_{\text{MSS}} = \begin{bmatrix} \mathbf{A}_{cs1} & \mathbf{A}_{cs2} & \mathbf{A}_{cs3} & \mathbf{A}_{FE} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs1} \\ \mathbf{x}_{cs2} \\ \mathbf{x}_{cs3} \\ \mathbf{x}_{FE} \end{bmatrix} . \quad (24)$$

268 Here, $\bar{\mathbf{l}}_{\text{MSS}} = \mathbf{l}_{\text{MSS}} - \Delta \mathbf{l}_{\text{MSS}}$ and $\bar{\Sigma}_{\text{MSS}} = \Sigma_{\text{MSS}} + \Sigma_{\Delta \text{MSS}}$ are different for
 269 each of the three previous approaches. The overall gravity field observation

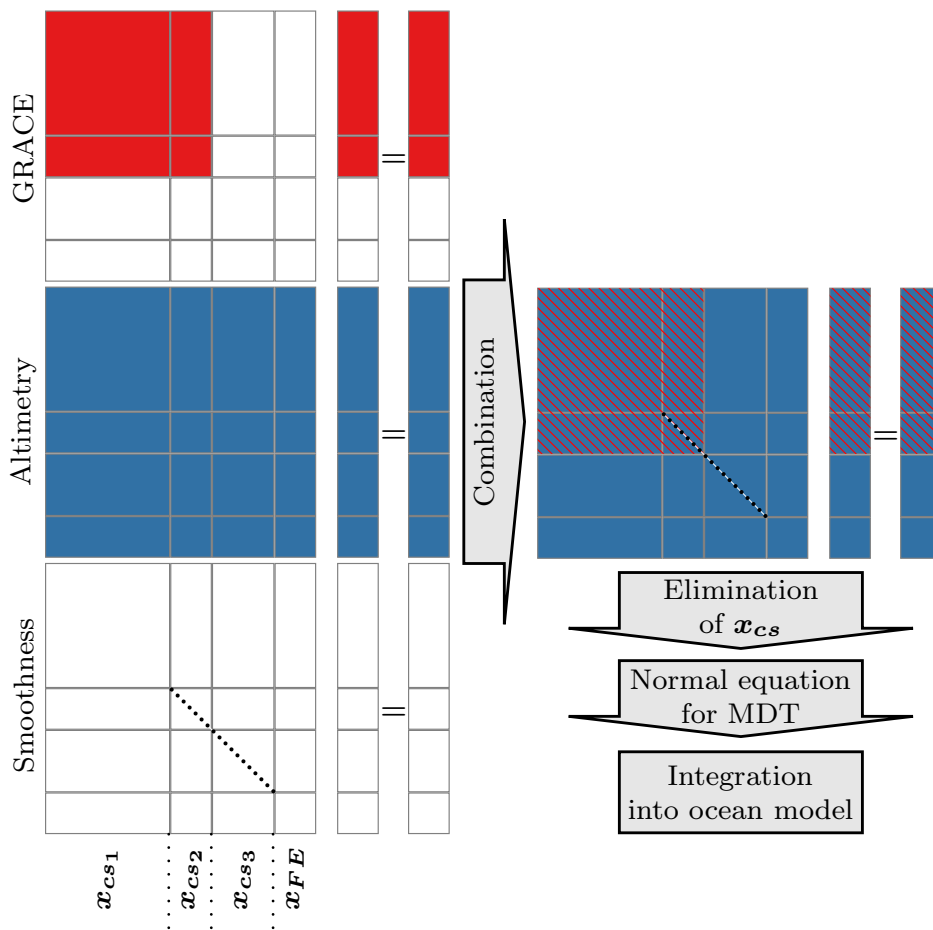


Figure 2: Schematic of the particular normal equations and the rigorous combination model

270 equations are

$$271 \quad \begin{bmatrix} \mathbf{x}_1^{\text{ITG}} \\ \mathbf{x}_2^{\text{ITG}} \end{bmatrix} + \begin{bmatrix} \mathbf{v}_1^{\text{GRACE}} \\ \mathbf{v}_2^{\text{GRACE}} \end{bmatrix} = \begin{bmatrix} \mathbf{I} & \mathbf{0} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{I} & \mathbf{0} & \mathbf{0} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs1} \\ \mathbf{x}_{cs2} \\ \mathbf{x}_{cs3} \\ \mathbf{x}_{FE} \end{bmatrix} \quad (25)$$

272 with the covariance matrix Σ_{ITG} . The pseudo-observation equations for the
273 introduced smoothness conditions are

$$274 \quad \begin{bmatrix} \mathbf{0} \\ \mathbf{0} \end{bmatrix} + \begin{bmatrix} \mathbf{v}_1^{\text{smooth}} \\ \mathbf{v}_2^{\text{smooth}} \end{bmatrix} = \begin{bmatrix} \mathbf{0} & \mathbf{1} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{1} & \mathbf{0} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs1} \\ \mathbf{x}_{cs2} \\ \mathbf{x}_{cs3} \\ \mathbf{x}_{FE} \end{bmatrix} \quad (26)$$

275 with the stochastic information contained in

$$276 \quad \Sigma_{cs}^{\text{smooth}} = \begin{bmatrix} \text{diag}(\sigma_{nm_2}^2) & \mathbf{0} \\ \mathbf{0} & \text{diag}(\sigma_{nm_3}^2) \end{bmatrix}. \quad (27)$$

277 In contrast to (14) which describes the complete observation equations, the
278 parameters \mathbf{x}_{cs4} are no longer required here. Because the \mathbf{x}_{cs4} are separated
279 from the vector of unknowns, additional smoothness conditions need not be
280 applied in the corresponding frequency domain.

281 Subsequently the normal equations for the particular groups of observations
282 can be accumulated. Figure 2 shows a schematic diagram of the resulting
283 normal equations. After renaming the coefficients and the right-hand side of
284 equation (8) these are written as

$$285 \quad \begin{bmatrix} \mathbf{N}_{cs} & \mathbf{N}_{cs,FE} \\ \mathbf{N}_{FE,cs} & \mathbf{N}_{FE} \end{bmatrix} \begin{bmatrix} \mathbf{x}_{cs} \\ \mathbf{x}_{FE} \end{bmatrix} = \begin{bmatrix} \mathbf{n}_{cs} \\ \mathbf{n}_{FE} \end{bmatrix}. \quad (28)$$

286 The gravity field parameters \mathbf{x}_{cs} can be eliminated by a Schur decomposition
 287 from these normal equations to provide the normal equations for the mean
 288 dynamic topography

$$\begin{aligned}
 & (\mathbf{N}_{FE} - \mathbf{N}_{FE,cs} \mathbf{N}_{cs}^{-1} \mathbf{N}_{cs,FE}) \mathbf{x}_{FE} = (\mathbf{n}_{FE} - \mathbf{N}_{FE,cs} \mathbf{N}_{cs}^{-1} \mathbf{n}_{cs}) \\
 & \Sigma_{FE}^{-1} \mathbf{x}_{FE} = \bar{\mathbf{n}}_{FE} \quad .
 \end{aligned}
 \tag{29}$$

290
 291 At this point the great advantage and benefit of the rigorous combination
 292 model becomes clear. As mentioned above the finite elements are directly
 293 applied to the nodal points of the ocean circulation model. Thus the resulting
 294 normal equation matrix represents the inverse covariance matrix Σ_{FE}^{-1} of the
 295 mean dynamic topography required by the ocean circulation model; that is,
 296 no additional inversion is required to compute a weight matrix from an error
 297 covariance matrix and the MDT can be directly combined with the ocean
 298 circulation model.

299 3. Numerical results

300 The static gravity field solution ITG-Grace2010s (Mayer-Gürr et al.,
 301 2010) and the altimetric mean sea surface model MSS_CNES_CLS10 (MSS_CNES_CLS10,
 302 2010) are used in this study. The MSS is given on a regular grid with a res-
 303 olution of $(1/30)^\circ$ covering the global oceans between the latitudes $80^\circ S$ and
 304 $84^\circ N$. MSS_CNES_CLS10 includes an error estimate Σ_{MSS} that mainly re-
 305 flects the formal errors of the optimal interpolation method used in mapping.
 306 Thus, the formal errors may not account for other errors such as radial or-
 307 bit errors. In this experiment it is assumed that these additional errors are
 308 very small compared to the stochastic characteristics of the random field \mathbf{S} ,

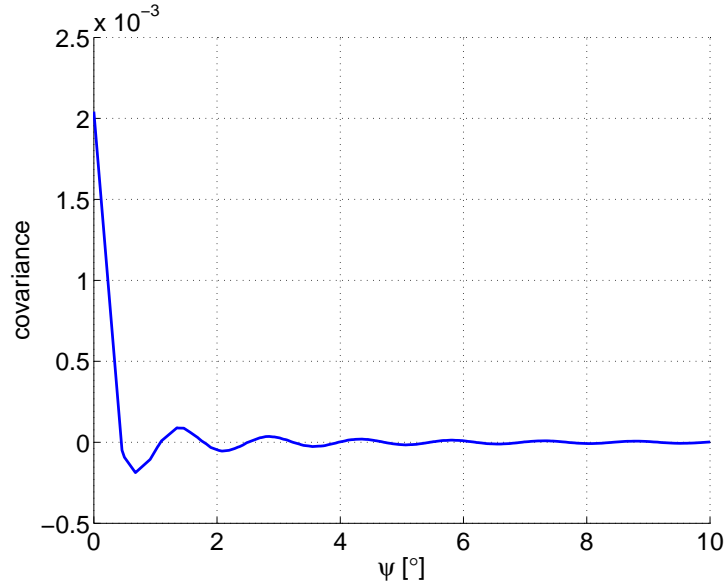
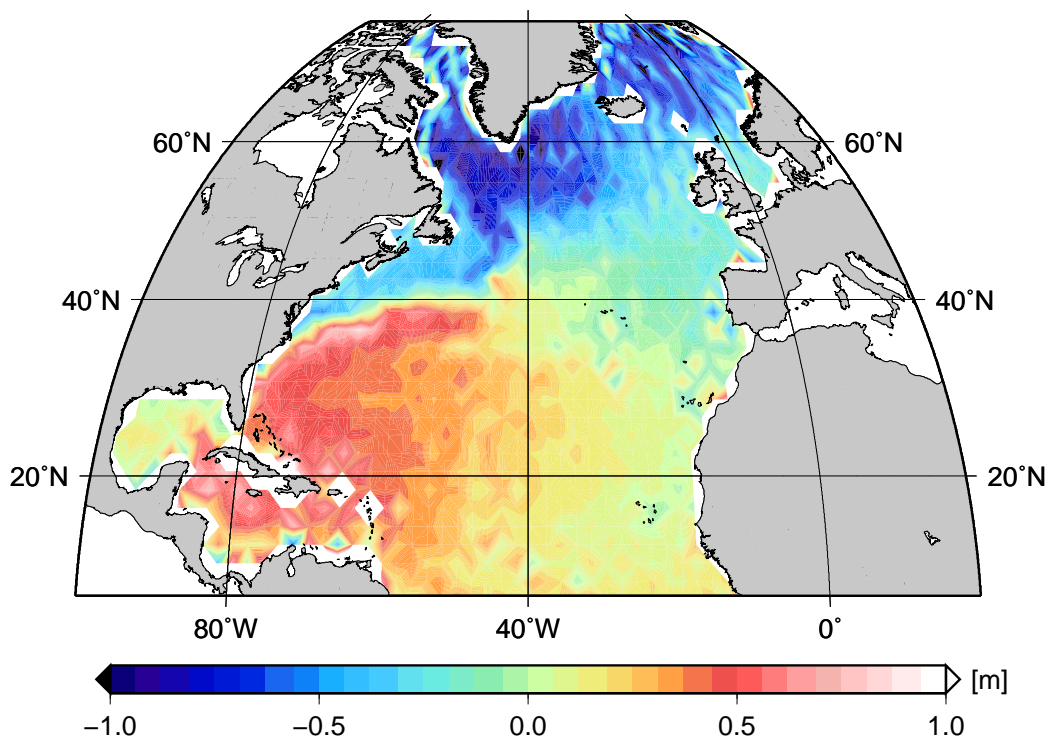


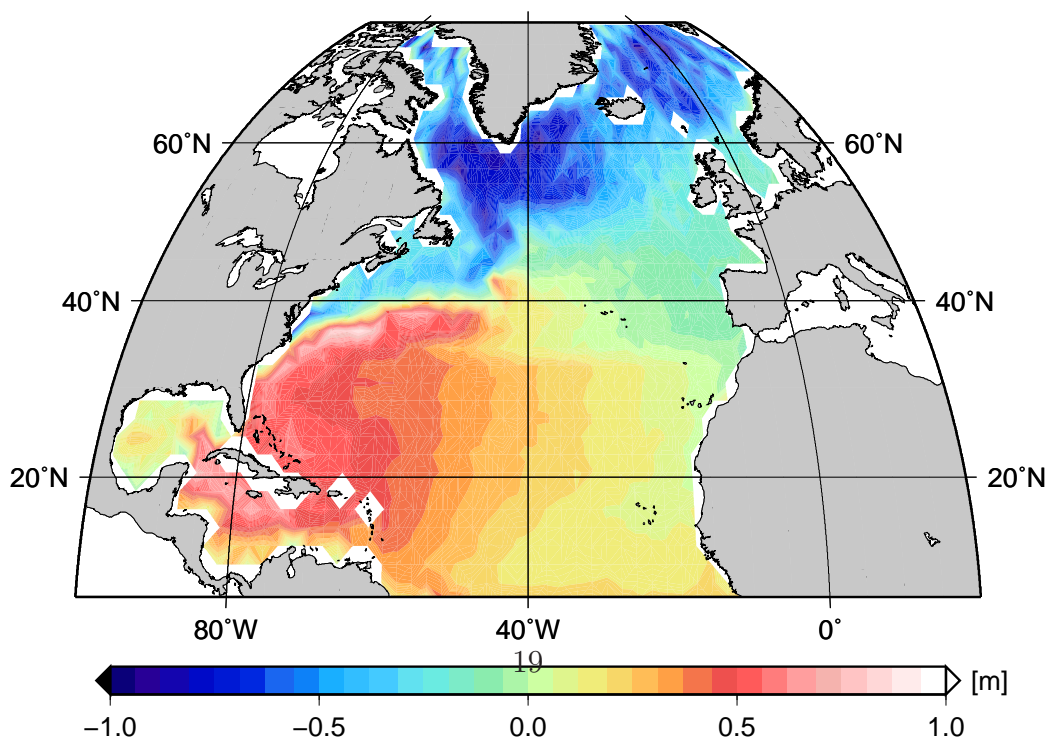
Figure 3: Covariance function for approach Rifugio03.

309 expressed by $\Sigma_{\Delta\text{MSS}}$ (cf. table 1). The MSS as well as the respective errors
310 for the North Atlantic Ocean are extracted from the original data set. To
311 reduce the computational effort only values on a $0.5^\circ \times 0.5^\circ$ grid are selected.
312 Figure 3 shows that for the covariance function following approach Rifugio03
313 (section 2.2.3, equation (22)) point values at a distance of 0.5° can be con-
314 sidered as nearly uncorrelated because the central maximum falls off very
315 quickly. This observation also holds for the covariance models obtained by
316 approaches Rifugio01 and Rifugio02. Thus, correlations of MSS errors are
317 neglected in this study and a diagonal covariance matrix $\Sigma_{\Delta\text{MSS}}$ is used so
318 that the overall covariance matrix for the MSS results in

$$319 \quad \bar{\Sigma}_{\text{MSS}} = \Sigma_{\text{MSS}} + \sigma^2 \mathbf{I} \quad . \quad (30)$$



(a)



(b)

Figure 4: Mean dynamic topographies for approach Rifugio01 (a) and Rifugio03 (b).

	$\sigma[\text{m}]$	$\sigma_{\text{total}}[\text{m}]$
Rifugio01	$\sigma_{\text{emp}} = \pm 0.385\text{m}$	$\pm 0.386\text{m}$
Rifugio02	$\sigma_{\text{Kaula}} = \pm 0.265\text{m}$	$\pm 0.266\text{m}$
Rifugio03	$\sqrt{\sigma_{\text{EGM08}}^2 + \sigma_{\text{Kaula}}^2} = \sqrt{\pm 0.034^2\text{m}^2 + \pm 0.029^2\text{m}^2} = \pm 0.045\text{m}$	$\pm 0.048\text{m}$

Table 1: Standard deviations obtained with different approaches and representative values for the overall standard deviations.

320 The errors of the MSS_CNES_CLS10 representing the Σ_{MSS} range between
321 0.3 cm and 9.69 cm. The resulting standard deviations σ obtained by the
322 different approaches as well as a representative value for the overall standard
323 deviation are listed in table 1.

324 The finite elements are used on a triangulated $2^\circ \times 2^\circ$ grid with continuous
325 linear polynomials as basis functions. No additional smoothness conditions
326 are applied. The definition of this coarse grid ensures that (29) is solvable
327 and an analysis and a comparison of different mean dynamic topographies is
328 possible in this study. As a prerequisite the spatial resolution of the finite ele-
329 ments has to cover the frequency range for that both the GRACE geoid model
330 and the altimetric mean sea surface provide information with high accuracy
331 (see figure 1). Figure 4 shows the mean dynamic topography for Rifugio01
332 and Rifugio03. As expected, the Rifugio01 solution is less smooth, probably
333 because the high frequencies in the altimetric measurements leak into the
334 solution. Because of the consistent treatment of signal and omission error,
335 however, the resulting standard deviations also increase with the decreased
336 smoothness of the solution, as shown in figure 5. For a section along longitude

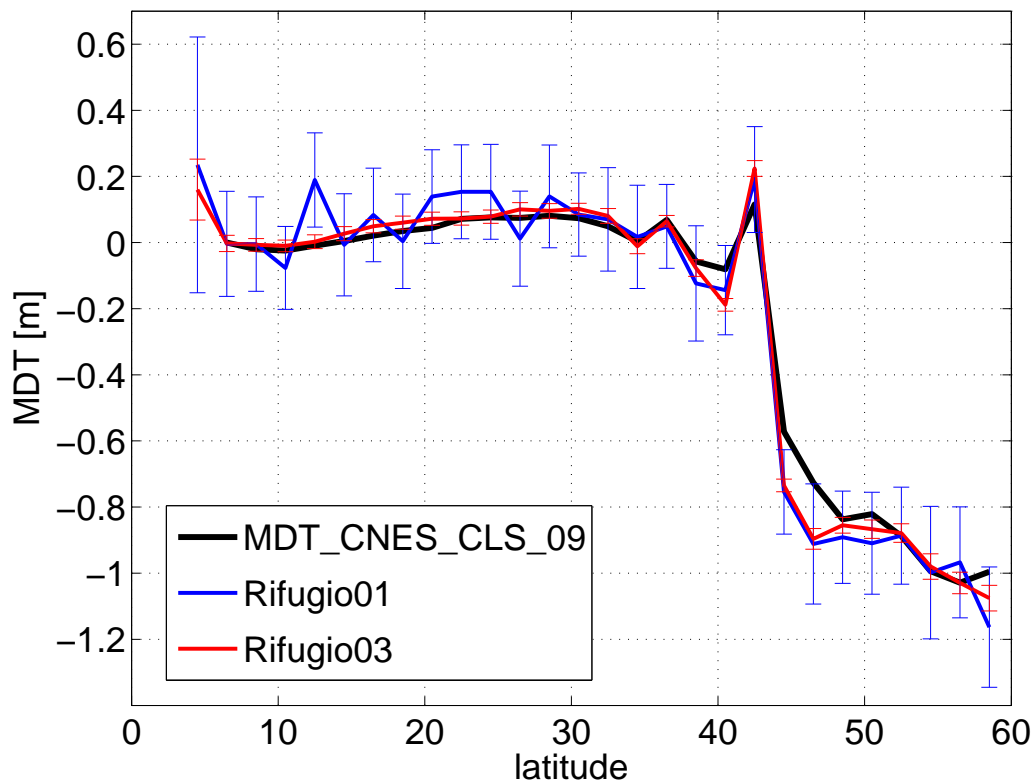


Figure 5: MDT along longitude -45.5° with error bars.

337 -45.5° , the mean dynamic topography agrees mostly within the correspond-
 338 ing error bars with the mean dynamic topography model MDT_CNES_CLS09
 339 (MDT_CNES_CLS09, 2009).

340 Figure 6 shows a histogram of differences between the MDT_CNES_CLS09
 341 estimate and Rifugio01 and Rifugio03 (gray bars). For Rifugio03 the distri-
 342 bution has a sharp peak near zero and the root-mean-square (rms) difference
 343 is 0.1072 m. For the Rifugio01 solution the distribution of differences is much
 344 broader (outliers, that are found mostly near the coast lines, are not shown
 345 for clarity) with an rms-difference of 0.2231 m.

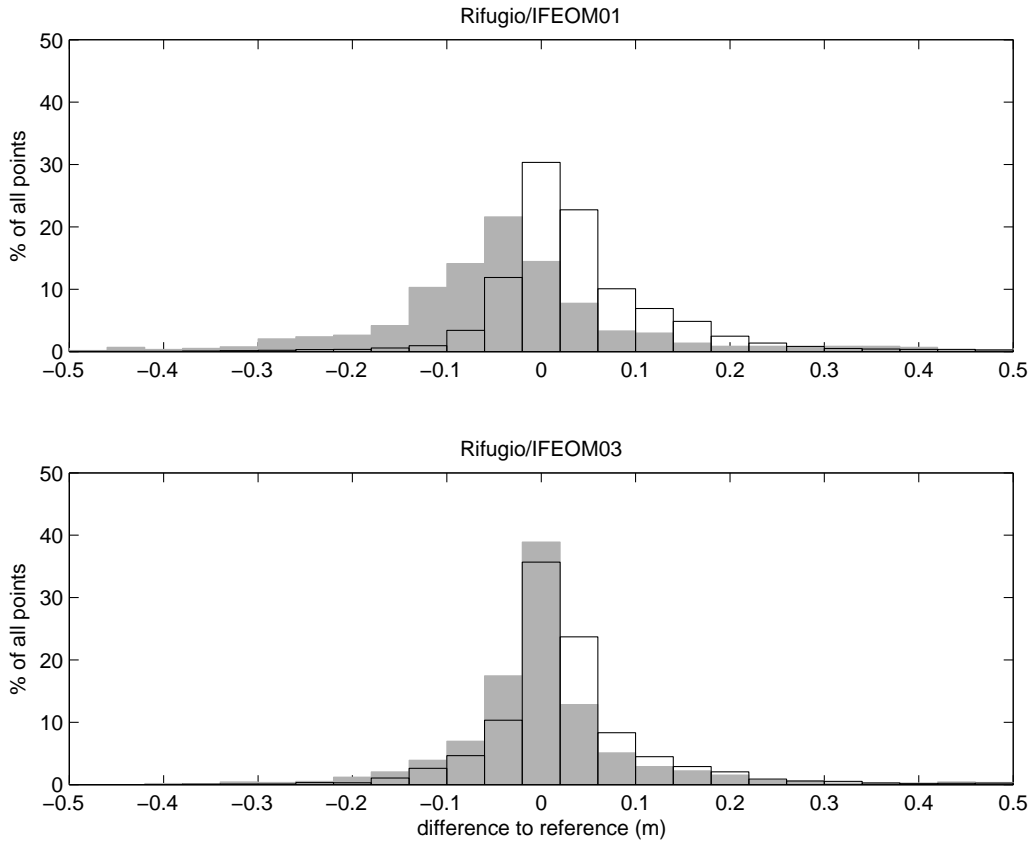


Figure 6: Histogram of difference to MDT estimate by CLS Space Oceanography Division. The gray bars show the difference between of Rifugio01 (top) and Rifugio03 (bottom) solutions to the MDT_CNES_CLS09 estimate (MDT_CNES_CLS09, 2009). The black lines indicate the differences of the corresponding IFEOM solutions after combination with the Rifugio estimates.

346 The effects of integrating the different solutions into the Inverse Finite
347 Element Ocean Model (IFEOM) are described in the following section.

348 4. Integration into IFEOM

349 4.1. Optimization by IFEOM

350 The MDT solutions Rifugio01 and Rifugio03 are combined with the In-
351 verse Finite Element Ocean Model (IFEOM) as described in section 2. The
352 results are labeled IFEOM01 and IFEOM03. The two MDT estimates repre-
353 sent the two extremes: for the estimate Rifugio01 minimal prior information
354 was assumed—the omitted signal has the expectation zero with a large em-
355 pirical error variance (section 2.2.1); in contrast, the estimate Rifugio03 was
356 obtained by assuming maximal prior information about the omission error—
357 the omitted signal is assumed to be estimated by EGM08 (section 2.2.3).
358 The resulting estimate is not only smoother for Rifugio03 than for Rifugio01
359 (as discussed in section 3, figures 4a and 4b), but also the error estimate
360 is much smaller for Rifugio03 (figure 5) so that the ocean model estimate
361 IFEOM03 is closer to Rifugio03 than IFEOM01 to Rifugio01.

362 Figure 6 shows that in spite of the large errors (small weights in the cost
363 function) in Rifugio01 the resulting IFEOM estimate (IFEOM01) matches
364 the MDT_CNES_CLS09 estimate better than the Rifugio01 estimate; the cor-
365 responding rms-difference is almost as small as that for Rifugio03: 0.1079 m.
366 The small errors (large weights in the cost function) of Rifugio03 make the
367 IFEOM03 solution adjust closely to Rifugio03 so that in the histogram of
368 differences to the MDT_CNES_CLS09 estimate there is only a small change
369 in the bias. The rms-difference is only slightly reduced to 0.0964 m. This

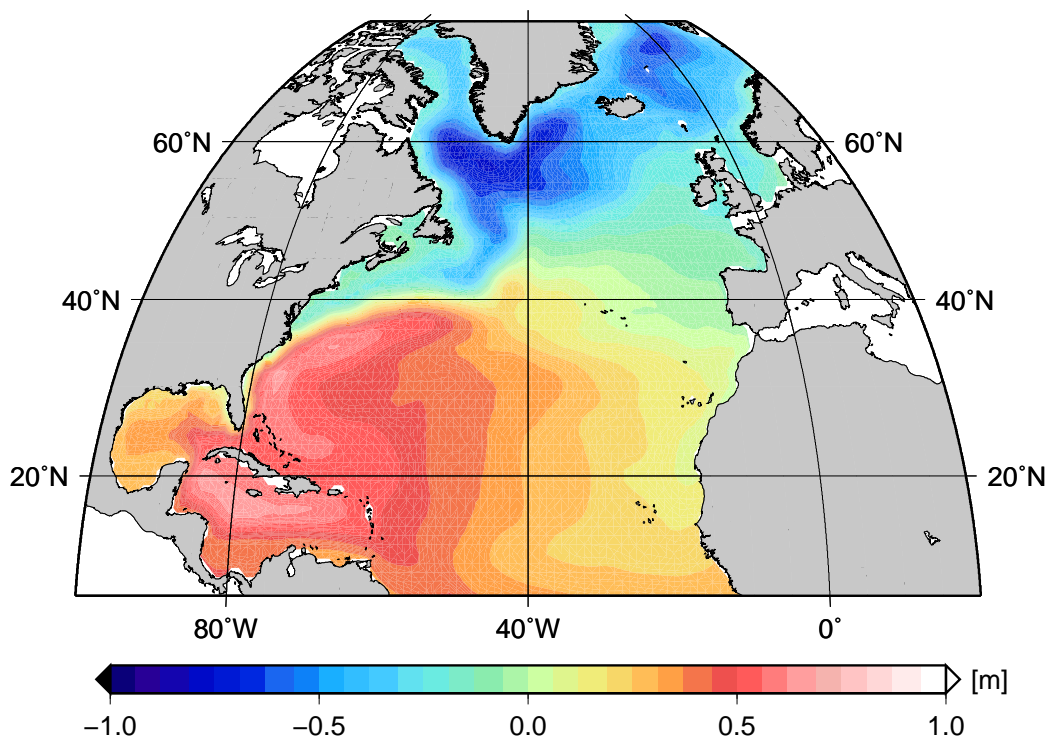
370 implies that IFEOM adds information (from other data sources such as hy-
371 drography) to the inaccurate estimate Rifugio01 to improve the MDT, while
372 IFEOM can barely change the more accurate estimate Rifugio03.

373 Figure 7 shows the two MDT estimates IFEOM01 and IFEOM03. The
374 optimization procedure rejects, based on the prior error estimates, the small
375 scale structures still apparent in figures 4a and 4b as unphysical noise, so
376 that both IFEOM estimates are smooth. This is interpreted as a success of
377 the consistent error description of Rifugio01 and Rifugio03.

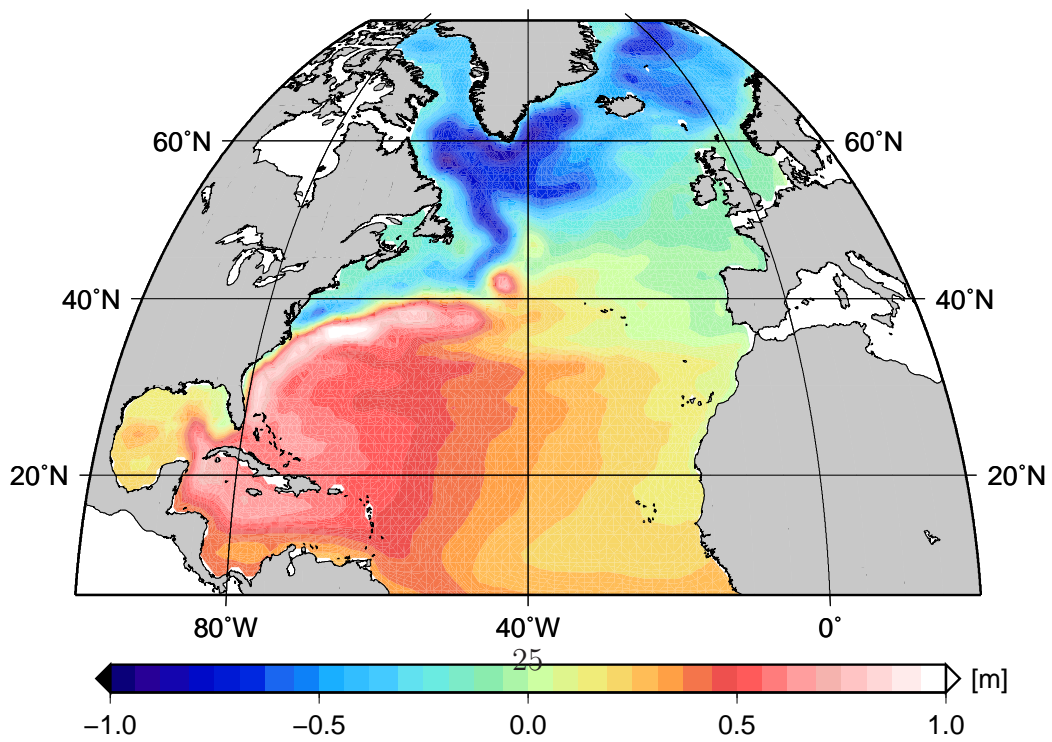
378 In the following the solutions IFEOM01 and IFEOM03 are compared
379 to previous solutions by Richter (2010). Richter obtained his solutions by
380 assimilating the Rio05 MDTRio and Hernandez (2004) and sea level anoma-
381 lies provided by Aviso (Archiving, Validation and Interpretation of Satellites
382 Oceanographic data, www.aviso.oceanobs.com). In the absence of any er-
383 ror estimation for these data, he used an annual variability (variance) in
384 the weighting procedure. This weighting approach is not comparable to the
385 methods described here, but Richter’s solutions still provide a well tuned
386 baseline for plausibility comparisons.

387 *4.2. Influence of the new data combination on oceanographic features*

388 The IFEOM03 solution has some remarkable new features that are a
389 consequence of both the new gravity field data and the new combination
390 method. Large differences in temperature compared to a solution of Richter
391 (2010) can be found in the Gulf Stream area. The temperatures at 120 m
392 depth are higher at the southern flank of the current and lower at the north-
393 ern side so that the across-stream temperature difference is increased by up
394 to 10 °C (figure 8a). In contrast, the corresponding salinity difference is de-

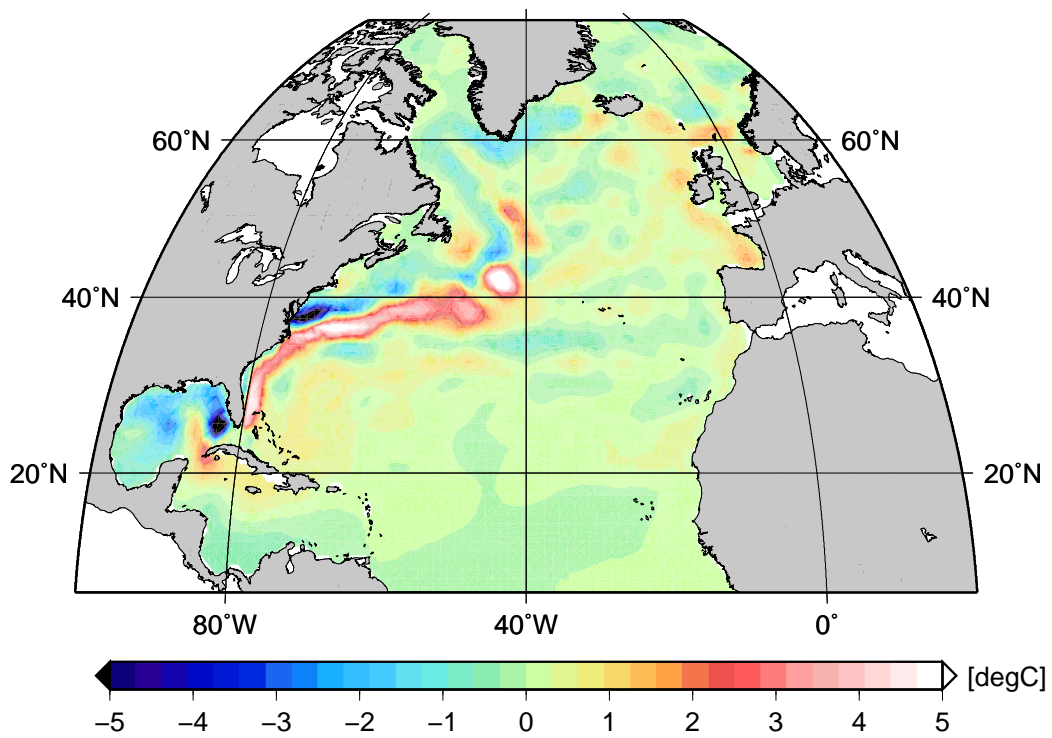


(a)

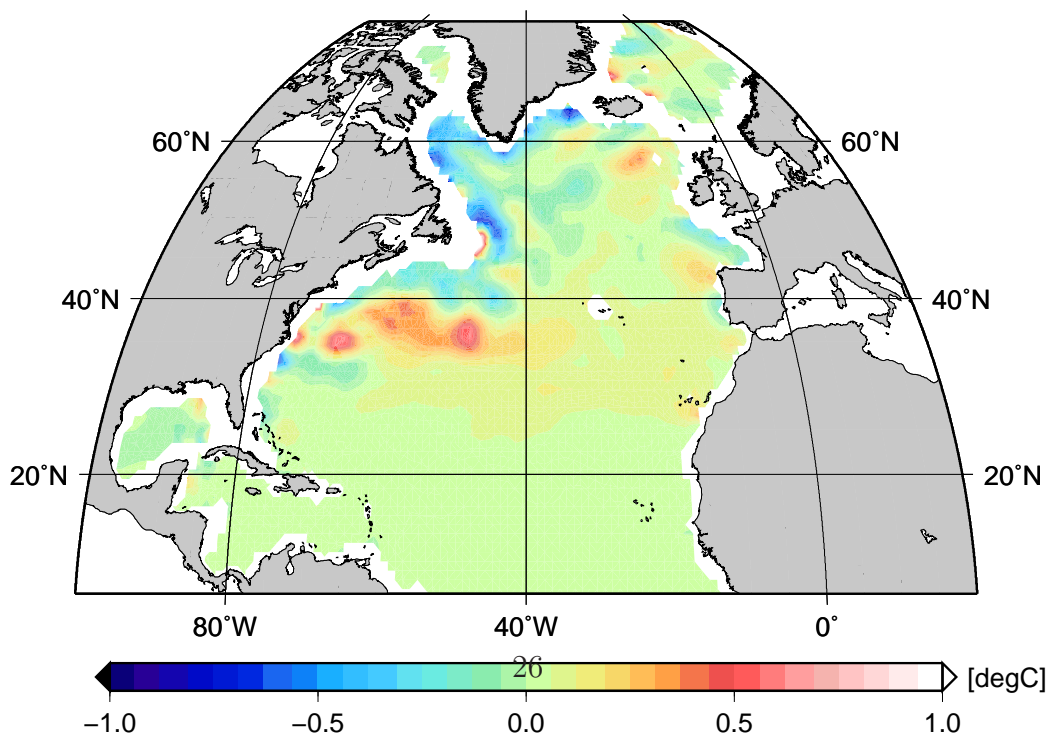


(b)

Figure 7: MDT estimates by IFEOM with Rifugio01: IFEOM01 (a) and Rifugio03: IFEOM03 (b).



(a)



(b)

Figure 8: Horizontal maps of temperature difference between IFEOM03 and Richter (2010) solutions at depths 120 m (a) and at 1000 m (b).

395 creased by about 1¹, so that, compared to the previous solution by Richter,
396 water is less saline at the southern boundary of the Gulf Stream (not shown).

397 At 1000 m depth, deep water masses along the coast of Greenland and in
398 the Labrador Sea are approximately 0.5 °C cooler (figure 8b) and 0.1 more
399 saline (not shown) in the IFEOM03 solution than in the Richter solution.
400 These differences are within the range of the assumed prior errors in this
401 model region of 2.48 °C and 0.35, respectively. This finding can be interpreted
402 as an increase in deep water formation rates when more cold and saline
403 surface water sinks to greater depths.

404 These characteristics are barely visible in the IFEOM01 solution (not
405 shown), because the large errors of Rifugio01 allow only small adjustments to
406 the MDT estimate thereby avoiding deviations from the first guess. The first
407 guess is a long term IFEOM model run on an extended model domain without
408 any satellite altimetry information, but only hydrographic data (Richter,
409 2010).

410 The meridional overturning stream function (figure 9) has changed un-
411 der the influence of the new MDT estimates. Both solutions show a new
412 maximum of over 20 Sv ($1 \text{ Sv} = 10^6 \text{ m}^3 \text{ s}^{-1}$) near 40–45°N that is missing
413 in Richter’s solution. The meridional circulation of the IFEOM01 solution
414 is weaker at the Southern boundary compared to Richter’s solution. This is
415 also apparent in the integrated meridional heat transports (figure 10). The
416 IFEOM03 solution is affected by the Rifugio03 MDT mostly in the Northern
417 part of the model area.

¹We use the practical salinity scale (PSS) for values of salinity. Note that in oceanogra-
phy, salinity is a conductivity ratio and therefore does not have units.

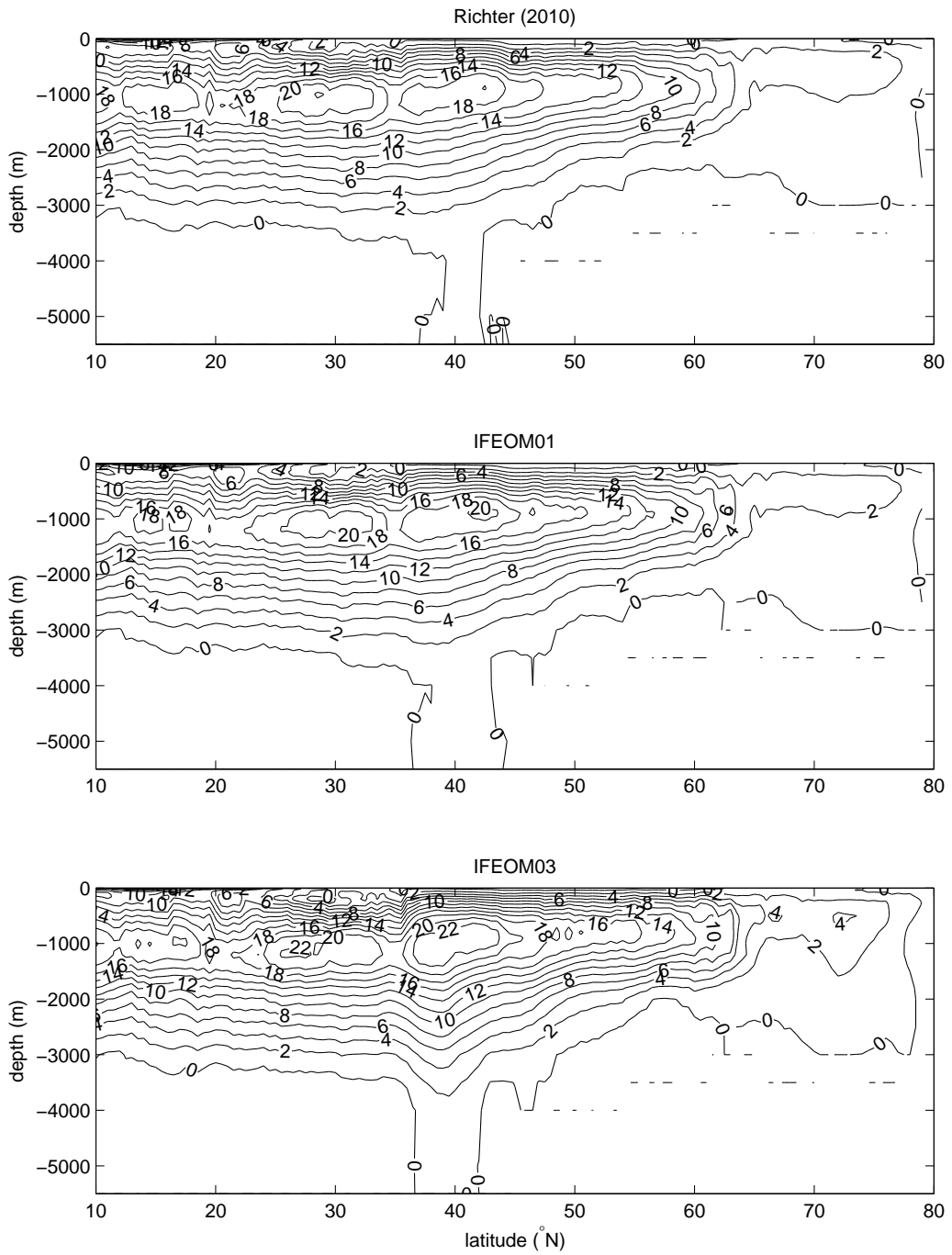


Figure 9: Overturning stream functions (in Sv).

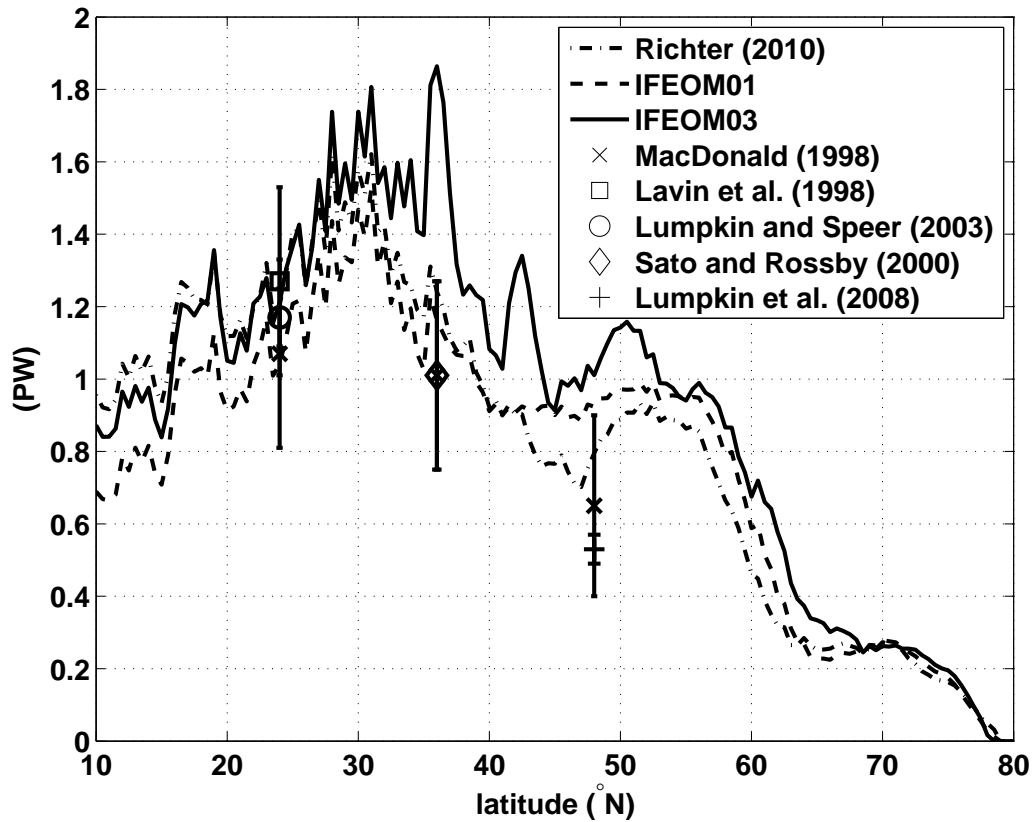


Figure 10: Heat transport estimates across latitudes (in PW), also included are previous estimates from individual section as listed in table 3.

418 Estimates of poleward oceanic heat transport differ for the various so-
 419 lutions. While Richter’s estimate is tuned to agree within error bars with
 420 almost all previous estimates in table 3 (and figure 10), the IFEOM03 solu-
 421 tion deviates from the estimates of Macdonald (1998) and Sato and Rossby
 422 (2000) for 36°N and from estimates of Lumpkin et al. (2008) and Macdon-
 423 ald (1998) for 48°N. Between latitudes of about 33° to 68°N, the IFEOM03
 424 solution transports more heat than established estimates (an incomplete list

Section	Heat transport in [PW]		
	Richter (2010)	IFEOM01	IFEOM03
24°N	1.20	1.04	1.21
36°N	1.24	1.16	1.86
48°N	0.80	0.92	1.01

Table 2: IFEOM heat transport estimates through zonal sections across the North Atlantic.

425 is found in table 3). There are two distinct peaks near 37° and 43°N that
426 are attributed to the strengthened circulation in the Gulf Stream region.
427 However, IFEOM03 estimates agree within error bars with previous results
428 for low latitudes up to 33°N, whereas IFEOM01 results are smaller for these
429 lower latitudes (as discussed above).

430 5. Conclusions

431 Estimates of the mean dynamic topography derived from satellite ob-
432 servations are useful for improving ocean circulation estimates, but only if
433 they have consistent error estimates. The procedure presented in this pa-
434 per achieves this goal by modeling the MDT on arbitrary ocean model grids
435 as the difference between altimetric sea surface and geoid height in a com-
436 bined estimation process. As a central feature of this process, the omission
437 error is treated explicitly. Different assumptions about the omission error
438 lead to MDT estimates that are different in resolved signal and estimated
439 error covariance. Assuming little prior knowledge about the omission error

Section	Heat transport in [PW] with errors	Source
24°N	1.07 ± 0.26	Macdonald (1998)
	1.27 ± 0.26	Lavin et al. (1998))
	1.17 ± 0.08	Lumpkin and Speer (2003)
36°N	1.01 ± 0.26	Macdonald (1998)
	1.2 ± 0.3	Sato and Rossby (2000)
48°N	0.65 ± 0.25	Macdonald (1998)
	0.53 ± 0.04	Lumpkin et al. (2008)

Table 3: Heat transport estimates of other authors through zonal sections across the North Atlantic.

440 leads to large uncertainties in the model MDT, while using the EGM08 as
441 the best available estimate of the omission error reduces these uncertainties
442 dramatically.

443 The design of the estimation process aims at using its products in inverse
444 problems in oceanography. For this purpose the error covariance matrix need
445 not be computed explicitly, but its inverse is used. The inverse error covari-
446 ance is, by design, given exactly by the normal equations of the estimation
447 problem.

448 The estimation process works for any target (ocean model) grid to give
449 consistent solutions for ocean modeling. The associated normal equations
450 can be solved exactly, when all observations resolve the grid-scale. For grids
451 finer than the resolution of the satellite observation products, the normal

452 equations are singular, but as the inverse of the normal equation matrix is
453 not required for the ocean model inversion, the method can also be applied.

454 The MDT estimates by the stationary inverse finite element ocean model
455 IFEOM are smooth even when the Rifugio estimate that is used in the inver-
456 sion is not. This implies that the short scales present in the Rifugio solutions
457 are unphysical, but they are successfully rejected by the ocean model because
458 of the consistent error estimates implicit in the normal equations. This should
459 be interpreted as the main piece of evidence that the assumptions, especially
460 about the omission errors, that went into the geodetic estimation process are
461 consistent with the ocean model. The ocean model IFEOM helps to improve
462 the Rifugio estimates of MDT. The less is assumed about the omission error,
463 the more the MDT estimates benefit from the ocean model contribution.

464 Fitting IFEOM to the Rifugio MDT generally accelerates the circula-
465 tion in the model ocean. The model result shows a more pronounced Gulf
466 Stream, increased deep water formation at high latitudes and modified merid-
467 ional heat transport estimates. Some of these estimates are not consistent
468 with previous estimates. These small discrepancies are attributed to tuning
469 issues in both the geodetic and the oceanographic estimation procedure and
470 possibly incomplete ocean model dynamics. In this context, stationarity (no
471 time dependence) appears as the main ocean model deficit that needs to be
472 addressed in the future.

473 Ocean modeling can greatly benefit from space-borne observations. Here,
474 the prospect of consistent satellite-based estimates of MDT with errors on
475 the ocean models grid is put forward in a pilot study.

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