1	Climatic and in-cave influences on $\delta^{18}O$ and $\delta^{13}C$ in a stalagmite from northeastern
2	India through the last deglaciation
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24 Abstract

25 Northeastern India experiences extraordinarily pronounced seasonal climate, 26 governed by the Indian Summer Monsoon (ISM). The vulnerability of this region to 27 floods and droughts calls for detailed and highly resolved paleoclimate 28 reconstructions in order to assess the recurrence rate and driving factors of ISM changes. We use stable oxygen and carbon isotope ratios (δ^{18} O and δ^{13} C) from 29 30 stalagmite MAW-6 from Mawmluh Cave to infer climate and environmental 31 conditions in northeastern India over the last deglaciation (16-6 kyr BP). We interpret stalagmite δ^{18} O as reflecting ISM strength, while δ^{13} C appears to be driven by local 32 33 hydroclimate conditions. Pronounced shifts in ISM strength over the deglaciation are 34 apparent from the δ^{18} O record, similarly to other records from monsoonal Asia. The 35 ISM is weaker during the late glacial and the Younger Dryas, and stronger during the 36 Bølling-Allerød and Holocene. Local conditions inferred from the δ^{13} C record appear 37 to have changed less substantially over time, possibly related to the masking effect of 38 changing precipitation seasonality. Time series analysis of the δ^{18} O record reveals 39 more chaotic conditions during the late glacial, and higher predictability during the

40 Holocene, likely related to the strengthening of the seasonal recurrence of the ISM41 with the onset of the Holocene.

42

43 INTRODUCTION

44 The Asian Monsoon (AM) is characterized by the seasonal reversal of circulation 45 between ocean and landmasses throughout South-East Asia, resulting in pronounced 46 hydroclimate seasonality in the affected regions. Stalagmite oxygen isotope ratio 47 $(\delta^{18}O)$ records from the AM realm have provided crucial information on past climate 48 conditions in this densely populated region (e.g., Wang et al., 2005; Cheng et al., 49 2016a; Eroglu et al., 2016). Pronounced glacial-interglacial variations in AM strength 50 are found to be strongly influenced by Northern Hemisphere Summer Insolation 51 (NHSI) (Cheng et al., 2009, 2016a; Kathayat et al., 2016), and closely related to 52 changes in the North Atlantic (Wang et al., 2001; Yuan et al., 2004). The AM is 53 therefore a highly dynamic system susceptible to external and internal forcings, 54 calling for precise paleoclimate reconstructions throughout monsoonal Asia to infer 55 on future developments under climate change scenarios.

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57 The majority of precisely dated high-resolution reconstructions of past glacial-58 interglacial AM variation stem from Chinese caves, providing unprecedented insight 59 in monsoonal dynamics over the past 640,000 years (Cheng et al., 2016a). 60 Information from the Indian subcontinent, particularly at high temporal resolution and 61 chronological precision, is still relatively scarce over these timescales (e.g., Sinha et 62 al., 2005; Govil and Divakar Naidu, 2011; Zhisheng et al., 2011; Menzel et al., 2014; 63 Kathayat et al., 2016). The Indian Summer Monsoon (ISM), the branch of the AM 64 that delivers moisture from the Arabian Sea and Indian Ocean to the Indian 65 subcontinent, as well as to the Arabian peninsula (Burns et al., 2002; Fleitmann et al., 66 2007), and China (Zhisheng et al., 2011; Baker et al., 2015), delivers ~80% of the 67 annual rainfall of these regions, and dominates their hydrological cycle (Sinha et al., 68 2007, Breitenbach et al. 2010) (Fig. 1A).

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Stalagmite δ^{18} O is a widely applied proxy for monsoonal strength in Asia, and is interpreted as reflecting the δ^{18} O of precipitation (Breitenbach et al. 2010, 2015; Pausata et al., 2011; Cheng et al., 2016b; Eroglu et al., 2016). As the δ^{18} O signal is

governed by a multitude of factors, their relative importance at a specific location 73 74 needs to be carefully assessed in order to correctly interpret paleoclimatic data. 75 Isotopic composition of the moisture source, transport length, and the amount of precipitation at the site can all affect precipitation δ^{18} O in monsoonal regions, and 76 77 taken together they provide information on monsoonal strength (Breitenbach et al., 78 2010; Baker et al., 2015; Eroglu et al., 2016). Cave monitoring efforts and 79 simultaneous study of different stalagmite geochemical proxies in parallel often allow 80 to more clearly determine the controls on local climate conditions, leading to more 81 accurate paleoclimate reconstructions (Baldini, 2010; Oster et al., 2012; Breitenbach 82 et al., 2015; Baldini et al., 2016; Cheng et al., 2016b). Stable carbon isotope ratios $(\delta^{13}C)$ are routinely measured together with $\delta^{18}O$, but have so far rarely been reported 83 84 for records from monsoonal Asia. This is partly due to the more complicated and site-85 specific interpretation of stalagmite δ^{13} C, which necessitates thorough understanding 86 of the local conditions; $\delta^{13}C$ can be influenced by vegetation composition (i.e., C_3 vs. C₄ plants), soil processes, open vs. closed conditions in the karst during carbonate 87 88 dissolution, and isotope fractionation in or above the cave (Fairchild and Baker, 89 2012). However, carefully evaluated stalagmite $\delta^{13}C$ time series can provide 90 important climate information to supplement and extend the interpretation of δ^{18} O 91 records, often resulting in a more in-depth understanding of past climate conditions 92 (Genty et al., 2003; Cosford et al., 2009; Ridley et al., 2015; Cheng et al., 2016b). 93 Particularly interesting is the difference in spatial scale between both proxies: while 94 δ^{18} O generally reflects large-scale atmospheric circulation processes (Breitenbach et al., 2010; Baker et al., 2015), δ^{13} C is a proxy for local processes and therefore more 95 96 sensitive to changes at local to regional level (Ridley et al., 2015; Cheng et al., 97 2016b).

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Here we present new sub-decadally resolved δ^{18} O and δ^{13} C data from a stalagmite from northeastern (NE) India that covers the interval of the last deglaciation and early Holocene (~16-6.5 ky BP). The last deglaciation was a period of rapid and substantial global climate change, driven by a ~3.5°C increase in global temperatures (Shakun et al., 2012). This resulted in large-scale reorganizations of circulation and weather patterns globally, with important repercussions in monsoonal Asia (Dykoski et al., 2005; Cheng et al., 2009; Ma et al., 2012). This analysis follows long-term studies of 106 the controls on precipitation δ^{18} O (Breitenbach et al., 2010), as well as detailed cave 107 microclimate monitoring schemes (Breitenbach et al., 2015), which allow us to 108 disentangle local and regional responses to climate change over the last deglaciation.

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110 GEOGRAPHICAL AND CLIMATOLOGICAL SETTING

111 Mawmluh Cave is located at 25°15'44"N, 91°52'54"E, 1320 m above sea level on the 112 Meghalaya plateau in NE India (Fig. 1). The cave developed in a Tertiary limestone 113 butte at the southern fringe of the plateau (Ghosh et al., 2005; Gebauer, 2008), and is 114 today mainly covered by grassland. Mean annual air temperature inside the cave 115 (18.5°C) is very similar to that recorded at the meteorological station Cherrapunji 116 (17.4°C) and in the nearby Mawmluh village (19.1°C).

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118 Hydroclimate in Meghalaya is extremely seasonal, with ~80% of annual precipitation 119 falling during the ISM season (June-October; Breitenbach et al., 2010). The 120 Meghalaya Plateau is the first morphological barrier for northward-moving moisture 121 from the Bay of Bengal (BoB), inducing intense orographic rainfall. Thus, Meghalaya 122 is a major water source for the Bangladesh plains, a region frequently flooded during 123 summer, e.g., in 1998, when ~60% of the country was inundated (Murata et al., 2008; 124 Webster, 2013). Despite having the highest rainfall amount in the world (Prokop and 125 Walanus, 2003), low retention capacity, due to the geological conditions on the 126 southern Meghalaya Plateau, results in frequent water shortage during the dry season 127 (November-May).

128

129 MATERIALS AND METHODS

130 Stalagmite MAW-6

131 The 21 cm long stalagmite MAW-6 was found broken in Mawmluh Cave in 2009 and 132 its original position is known only approximately (Fig. 1). The stalagmite displays 133 complex brown-grey color variations, with bands up to a few mm wide, but no annual 134 laminations. At least three white layers can be discerned which span from the growth 135 axis towards the sides of the stalagmite. To verify the mineralogy in MAW-6, three 136 samples were analyzed by X-ray diffraction (XRD) using a powder XRD 137 diffractometer (Bruker, D8 Advance), equipped with a scintillation counter and an 138 automatic sampler at ETH Zurich, Switzerland.

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140 U-series dating and chronology development

141 After cutting the stalagmite lengthwise using a diamond stone saw, 24 U-series 142 samples with weights between 88 and 311 mg were milled using an ethanol-cleaned 143 stainless steel bit, and subsequently analyzed by multi-collector inductively coupled 144 plasma mass spectrometry using a Thermo-Finnigan Neptune in the Minnesota 145 isotope laboratory, University of Minnesota. The chemical procedures used to 146 separate uranium and thorium for U-series dating are similar to those described in 147 Edwards et al. (1987). Uranium and thorium isotopes were analyzed on the multiplier 148 behind the retarding potential quadrupole in peak-jumping mode. Instrumental mass 149 fractionation was determined by measurements of a ²³³U/²³⁶U spike. The detail techniques are similar to those described in (Cheng et al., 2000, 2009) and half-life 150 151 values are those reported in (Cheng et al., 2013).

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153 The age model for MAW-6 was computed for each growth segment by applying a 154 cubic interpolation procedure using the COPRA software (Breitenbach et al., 2012). 155 COPRA computed 2000 ensemble realizations for both the δ^{18} O and δ^{13} C records, 156 from which the median, i.e., the central age for a defined sample depth, was 157 calculated. The uncertainty in the age model is defined by the 95% confidence 158 intervals, derived using the ±2 σ deviation from the median (Breitenbach et al., 2012).

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160 Stable isotope analysis

161 1050 samples for stable isotope analysis were milled continuously at 200 μm 162 resolution using a semi-automated high-precision drill (Sherline 5400 Deluxe) at ETH 163 Zurich. Nine Hendy tests were performed over the length of the stalagmite to look for 164 signs of kinetic isotope fractionation effects and, if present, to evaluate potential 165 changes in the intensity of kinetic fractionation through time (Suppl. fig. 1). For this, 166 carbonate samples were drilled point-wise along a single layer of the stalagmite using 167 a 0.3 mm diameter drill bit.

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169 Samples were analyzed for δ^{18} O and δ^{13} C on a Delta V Plus mass spectrometer 170 coupled to a ThermoFinnigan GasBench II carbonate preparation device at the 171 Geological Institute, ETH Zurich (Breitenbach and Bernasconi, 2011). An in-house

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172 carbonate standard (MS2), which is well linked to NBS19 (Breitenbach and 173 Bernasconi, 2011), was used to evaluate the runs. All values are expressed in permil 174 (‰) and referenced to the Vienna PeeDee Belemnite (V-PDB) standard. The external 175 standard deviation (1 σ) for both, δ^{18} O and δ^{13} C analyses on the carbonate is smaller 176 than 0.07‰.

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178 Because the MAW-6 record covers the period of the last deglaciation, the contribution 179 of changes in both sea surface temperature (SST) and sea level due to the melting of continental ice sheets to stalagmite δ^{18} O must be considered. To estimate this 180 contribution, a linear interpolation of seawater δ^{18} O values (δ^{18} O_{seawater}) reconstructed 181 from a sediment core from the BoB (Rashid et al., 2011) was performed to fit the 182 183 MAW-6 data points. The $\delta^{18}O_{\text{seawater}}$ record was subsequently subtracted from the 184 measured $\delta^{18}O_{calcite}$ in MAW-6, to yield an ice volume corrected ($\delta^{18}O_{IVC}$) record 185 (Suppl. Fig. 2). It should be noted that this procedure can introduce artefacts, as the 186 records are irregularly sampled, and the results should be interpreted with care.

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188 Recurrence quantification analysis

189 To infer possible changes in the dynamical regime of the ISM between late glacial 190 and early Holocene conditions, we performed a statistical analysis considering the 191 deterministic nature of the underlying process (the ISM), encoded by the recurrence properties of the $\delta^{18}O_{IVC}$ record. We use a measure of complexity, called *recurrence* 192 193 determinism (DET), which is derived from a recurrence plot, a graphical, binary 194 representation of pairs of times of similar values (actually states) within the time 195 series (see Supplemental information for further details; Marwan et al., 2007; Ozken 196 et al., 2015; Eroglu et al., 2016). DET reveals high values for deterministic processes 197 and regular (e.g., cyclic, periodic) variations, whereas more stochastic (i.e., random) 198 dynamics lead to low DET values. Moreover, the recurrence analysis is combined 199 with the pre-processing TACTS technique that allows detrending regularization of 200 irregularly sampled time series (see Supplemental information; Ozken et al., 2015).

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The $\delta^{18}O_{IVC}$ record was divided into two periods, the late glacial (LG, 16-13 kyr BP) and the early Holocene (EH, 9-6.5 kyr BP) and the DET measure were calculated for

both periods separately (Fig. 8). A statistical test based on a bootstrap approach was

205 performed to evaluate the significance of the variations in the DET measure. This test 206 provides a cumulative probability distribution of DET measures corresponding to the 207 null-hypothesis that there is no change in the dynamics of the underlying climate

process. From this test distribution the upper 95% confidence limit can be defined.

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- 209

210 **RESULTS**

211 Petrography and Mineralogy

212 XRD analysis reveals that stalagmite MAW-6 consists of calcite. The white layers 213 described above were identified as dirt layers in the stalagmite. They are well visible 214 at the fringes of the respective layers, while the stalagmite tip has been washed clean 215 by impinging water droplets (Fig. 1C). The deposition of silty material on the 216 stalagmite surface at these depths might indicate burial of MAW-6 by sediment re-217 deposition in the cave, inhibiting further growth. Several buried stalagmites have been 218 located in the cave (Suppl. Fig. 3) and sediment migration within the cave passage 219 appears to be an important process during high-discharge events of the cave stream. 220 However, caution must be applied with this interpretation, since dirt layers can also 221 originate from other processes, such as aerosol and dust deposition.

222

223 Age model

MAW-6 grew between ~16 and 6.5 kyr BP. The age model is based on 20 U-series dates, with analytical errors between ± 16 and ± 264 years (Fig. 2 and Table 1). Four dating samples contained high amounts of detrital thorium and were excluded from the final age model (shown in red in Fig. 2).

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Three hiatuses were identified in the depth-age relationship, coinciding with the white dirt layers in the stalagmite. User-specified hiatus depths of 69.46 mm, 111.86 mm and 147.26 mm from the stalagmite top allowed COPRA to split the age model construction into independent age models (before and after the hiatuses respectively). This procedure yielded a segmented depth-age chronology for the stalagmite, with hiatuses at 10.4-9.6 kyr BP, 11.6-10.8 kyr BP and 13-12.4 kyr BP. The details for the age modeling procedure can be found in Breitenbach et al. (2012).

Using the COPRA procedure, the age uncertainties of the MAW-6 record can betransferred from the age to the proxy domain (Breitenbach et al., 2012), which results

238 in a 95% confidence interval of possible proxy values at a given point in time. As a 239 consequence, it is not possible to determine the high-frequency variations within the 240 bounds of the confidence interval (Fig. 3A). Comparative discussions to other 241 paleoclimate records from the AM realm are therefore restricted to the median proxy 242 values in MAW-6 derived from the COPRA Monte Carlo modeling and to long-term 243 centennial changes. We still show the original MAW-6 isotope data for a tentative 244 comparison with other available records in Fig. 7, as this kind of uncertainty is 245 common to all paleoclimate records.

- 246
- 247 The MAW-6 δ^{18} O and δ^{13} C records

The δ^{18} O profile varies from -8.3% to -2.8% (Fig. 3A), with heavier values found in 248 249 the oldest part of the record (end of the last glacial), and lighter values in the youngest 250 part (the Holocene). Both, the Bølling-Allerød (B-A) interstadial, beginning at ~14.5 251 kyr BP with -1.5% shift, and the Younger Dryas (YD), between 12.6 and 11.6 kyr BP 252 and featuring the heaviest values of the entire record (~-3.5%), are clearly 253 demarcated. However, the exact beginning and the end of the YD in MAW-6 cannot 254 be defined, since the interval is bracketed by two hiatuses. The 850-year long hiatus 255 that masks the end of the YD is followed by a substantial 4.5% decrease in δ^{18} O, 256 marking the transition into the Holocene (~9.6 kyr BP). This decrease occurs in two 257 rapid stages, characterized by hiatuses, interrupted by an interval (~ 10.8 - 10.2 kyr BP) of relatively constant intermediate values (~-6.5%). Lowest δ^{18} O values (~-8%) 258 259 are found during the early Holocene (~9.6 kyr BP), slightly increasing towards the youngest part of the record. Our sea-level corrected $\delta^{18}O_{IVC}$ record shows that ice 260 261 volume and SST changes affect the isotope signature mainly before the YD, with only 262 minor impacts during the Holocene, accounting for $\sim 1/4$ (1‰) of the shift between 263 deglaciation and Holocene (Fig. 3A). We are therefore confident that the larger part 264 (3‰) of the variation in δ^{18} O during the deglaciation is attributable to changes in ISM 265 strength. For the following discussion, only the $\delta^{18}O_{IVC}$ record is considered.

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267 Compared to the δ^{18} O profile, the δ^{13} C profile is much more uniform, with variations 268 ranging between -1.2 and -6.6‰ and without clear trends over time. The heaviest 269 values are found during the early part of the record (late glacial-B-A, average -4‰).

270 In contrast to δ^{18} O, the YD period is characterized by slightly lighter values than

during the late glacial-B-A section (average -4.3%), whereas the transition into the Holocene leads to the most negative values (average -5.2% post ~9.6 kyr BP). The high-frequency variations in δ^{18} O and δ^{13} C are remarkably similar, but shifts in δ^{13} C are generally much more pronounced than in δ^{18} O. The similarity between the two records is also reflected by their high correlation (during all periods *r* > 0.55, Fig. 3B).

A crossplot of δ^{13} C vs. δ^{18} O_{IVC} reveals four distinct clusters (Fig. 4). These clusters 277 are mainly influenced by the average $\delta^{\rm 18}O_{\rm IVC}$ during the different periods, therefore 278 279 we distinguish a Holocene, intermediate (10.8-10.2 kyr BP), YD, and B-A/late glacial 280 cluster. The boxplot representation of the datasets in Fig. 4B allows the quantification 281 of temporal and proxy-related differences. While a clear distinction of the different 282 time periods is apparent in the $\delta^{18}O_{IVC}$ dataset, a much larger spread in $\delta^{13}C$ values is 283 found (Fig. 4B). The YD cluster is characterized by the heaviest $\delta^{18}O_{IVC}$ values of the entire record, while δ^{13} C is slightly lighter than during the late glacial. A trend 284 towards progressively lighter $\delta^{18}O_{IVC}$ values is found between the YD, intermediate, 285 and Holocene clusters, while intermediate δ^{13} C values are slightly heavier than during 286 287 the YD (Fig. 4B).

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The Hendy tests carried out throughout MAW-6 show evidence for kinetic fractionation during the YD and late glacial, while (near-)equilibrium conditions seem to have prevailed during the B-A and the Holocene (Suppl. Fig. 1). Kinetic effects are identified by strong correlations between δ^{18} O and δ^{13} C, as well as enrichment in the heavy isotopes with increasing distance from the growth axis (Hendy, 1971).

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295 Determinism of the δ^{18} O record

The analysis reveals distinctly different DET measures for the LG (DET=0.663 and inside the confidence interval) and the EH (DET=0.736, and outside the confidence interval) (Fig. 8). A high DET measure indicates a more predictable, i.e., a less chaotic, regime, while the opposite holds true for low DET measures.

300

301 **DISCUSSION**

302 The influence of karst processes on stalagmite stable isotopes in Mawmluh Cave

303 Karst processes at Mawmluh Cave are driven by the seasonal cycle in regional hydrology (Fig. 5). Precipitation δ^{18} O becomes increasingly lighter during the ISM 304 305 months, and reaches the most negative values during the late and post ISM (August-306 October) (Breitenbach et al., 2010). A direct amount effect can therefore be ruled out, 307 as maximum precipitation occurs earlier in the ISM season (July-August). Instead, 308 precipitation δ^{18} O in Meghalaya is controlled by: i) the travel distance of the air 309 masses, which increases throughout the ISM season, promoting stronger Rayleigh 310 fractionation during transport and lighter δ^{18} O at the site (Breitenbach et al., 2010), ii) 311 stronger contribution from isotopically depleted freshwater delivered to the BoB 312 during the late ISM (Sengupta and Sarkar, 2006; Singh et al., 2007; Breitenbach et al., 313 2010), and iii) isotopic depletion of rainwater during large rainstorms (amount effect 314 sensu Dansgaard, 1964) (Lawrence et al., 2004; Breitenbach et al., 2010; Baker et al., 315 2015). These mechanisms all drive precipitation δ^{18} O in the same direction, resulting 316 in lighter δ^{18} O during and after the ISM, and heavier values during dry season months 317 (Breitenbach et al., 2015; Myers et al., 2015). At Mawmluh Cave, infiltration is strongly skewed towards the summer months, and consequently dripwater δ^{18} O is 318 319 biased towards the ISM season (Fig. 5). Still, a clear seasonal cycle in dripwater δ^{18} O 320 is observed, with the lightest values occurring during the late ISM months, indicating 321 rapid (<1 month) fluid transfer into the cave (Breitenbach et al., 2015) (Fig. 5). 322 Dripwater (and stalagmite) δ^{18} O at Mawmluh Cave can therefore be used as a reliable 323 ISM strength proxy.

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325 Dripwater δ^{13} C can be influenced by changes in vegetation type above the cave (C₃ 326 vs. C₄ plants; Denniston et al., 2001), soil activity (Genty et al., 2006; Scholz et al., 327 2012), bedrock dissolution and open vs. closed system conditions in the karst (Genty 328 et al., 2001), and prior calcite precipitation (PCP) and fractionation processes in the 329 cave (Griffiths et al., 2012; Ridley et al., 2015). At Mawmluh Cave, precipitation and 330 consequently vegetation and soil activity (microbial activity and root respiration) are 331 at a maximum during the summer months (June-October), resulting in highest relative 332 humidity and soil pCO₂ during this period. The extremely high amounts of rainfall 333 delivered at Mawmluh Cave during the ISM season (max. 13,472 mm between June 334 and September; Breitenbach et al., 2015) lead to waterlogging of the soil and karst 335 overlying the cave (Breitenbach et al., 2015), most likely resulting in more closed

336 system conditions. Therefore, prior carbonate precipitation (PCP) in epikarst and cave 337 is minimized (or even completely absent) during the ISM season. Strong seasonal 338 variations in cave air pCO₂ are observed as a consequence of seasonal ventilation 339 changes (Breitenbach et al., 2015) (Fig. 5). During the dry season months, low cave 340 air pCO₂ due to strongly reduced rainfall amount above the cave and intensified 341 ventilation, leads to enhanced degassing of CO₂ from the solution, enriching 342 dripwater in ¹³C (Breitenbach et al., 2015). Moreover, open system conditions prevail 343 in the overlying soil and karst, due to seasonal aridity, resulting in low soil activity 344 (less input of isotopically light organic carbon to soil water) and promoting PCP (Fig. 345 5). All factors taken together, conditions during the dry season result in heavier dripwater δ^{13} C in Mawmluh Cave, whereas the opposite holds true for the ISM 346 347 months. Dripwater and stalagmite δ^{13} C is therefore strongly influenced by effective 348 infiltration in the soil and tightly connected to local climate conditions.

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350 Interpretation of the MAW-6 isotope records

We find large variations in $\delta^{18}O_{IVC}$ in stalagmite MAW-6 over the period of the last 351 352 deglaciation, with the heaviest values recorded during the late glacial and YD (Fig. 3). 353 Considering the controls on precipitation and dripwater $\delta^{18}O_{IVC}$ at Mawmluh Cave, we 354 interpret the late glacial and YD portions of the record as periods of weaker/shorter 355 ISM, accompanied by changes in the circulation regime (i.e., a more proximal 356 moisture source), whereas stronger ISM and longer moisture transport paths prevailed 357 during the B-A and the Holocene. Changes in both sea surface temperature (SST) and 358 sea level due to the melting of the continental ice sheets during the last deglaciation 359 resulted in substantial alteration of the isotopic composition of the surface ocean 360 water, as well as affecting evaporation and convection from the sea surface (Gadgil, 361 2003). The changes in the moisture source affect precipitation and stalagmite δ^{18} O. In 362 the BoB, the moisture source for the ISM, a ~3.2-3.5°C increase in SST between the 363 Last Glacial Maximum and the Holocene, and a +1.4°C SST shift between the YD 364 and the Holocene, have been documented (Rashid et al., 2007, 2011; Govil et al., 2011). In this region, additional depletion of seawater ¹⁸O occurred most likely due to 365 366 freshening of the BoB by increased runoff from precipitation and glacier melt in the 367 Himalaya and Tibet.

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The millennial-scale average in MAW-6 δ^{13} C shows much lower variability than 369 $\delta^{18}O_{IVC}$ over the last deglaciation, but the centennial-scale variations are remarkably 370 371 similar (Fig. 3). It is likely that changes in vegetation density and composition 372 occurred between cold/dry glacial and warm/humid interglacial periods. However, 373 vegetation changes above the cave as the primary cause for the high frequency 374 variation in δ^{13} C can probably be ruled out, as these would require longer time 375 periods and would likely be more gradual than the rapid decadal-scale shifts we find 376 in MAW-6. Karst processes, namely PCP and kinetic fractionation in the cave, can 377 best explain the observed variation in MAW-6 δ^{13} C. We find heavier δ^{13} C values during weak ISM periods, as identified in the $\delta^{18}O_{\rm IVC}$ record, indicating enhanced PCP 378 379 and kinetic fractionation stemming from drier summer and/or longer winter seasons. 380 Periods of strong ISM, on the other hand, are characterized by lighter δ^{13} C values, 381 which is in line with more closed system conditions during the wet summers, higher 382 cave air pCO₂, which subdues kinetic fractionation, and more active vegetation and 383 soil.

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It is possible that kinetic processes affect stalagmite δ^{18} O as well, precluding 385 386 quantitative rainfall reconstructions, but still allowing qualitative interpretation of monsoon strength. In fact, kinetic fractionation would drive stalagmite δ^{18} O towards 387 more positive values, as prolonged degassing and possibly evaporation enrich the 388 389 precipitating solution in the heavy isotope, thus increasing the sensitivity of the 390 speleothem to record dry periods. Modern dripwater δ^{18} O values directly reflect 391 precipitation δ^{18} O values at the site, lending additional confidence to the interpretation 392 of stalagmite δ^{18} O as a monsoon strength proxy. Periods of enhanced kinetic 393 fractionation in the past can be detected using the Hendy tests. Evidence for kinetic 394 fractionation is observed during the YD and the late glacial, periods that we interpret 395 as drier, while (near-)equilibrium conditions seem to have prevailed during the B-A 396 and the Holocene, when conditions were wetter (Suppl. Fig. 1). These results have to 397 be interpreted with care, however, as sampling along a single growth layer is 398 extremely difficult when no annual laminae are present.

399

400 We can use the complementary information of $\delta^{18}O_{IVC}$ and $\delta^{13}C$ in stalagmite MAW-6 401 to interpret climate variations on supra-regional as well as local scale over the last

deglaciation. Clear shifts in average $\delta^{18}O_{IVC}$ are apparent during different time periods 402 403 (clusters in Fig. 4), indicating changing ISM strength, related to the moisture source 404 and composition upstream of the study site. While shifts in δ^{13} C are less strongly 405 expressed, it is still possible to distinguish periods of local aridity/humidity related to 406 the amount of effective infiltration in the karst and cave ventilation dynamics. Positive correlation between $\delta^{18}O_{IVC}$ and $\delta^{13}C$ indicates that in general, weaker ISM 407 408 conditions are reflected as locally drier conditions at the study site, either due to 409 reduced summer rainfall, or a prolonged dry season (Fig. 4). "Weak-ISM" periods (late glacial, YD) are also characterized by a tendency towards heavier δ^{13} C, 410 suggesting drier conditions at the cave site, whereas Holocene $\delta^{18}O_{IVC}$ and $\delta^{13}C$ 411 clearly cluster at lighter values for both proxies, indicating strong ISM and humid 412 413 conditions at the study site.

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However, a detailed analysis of the relationship between $\delta^{18}O_{IVC}$ and $\delta^{13}C$ suggests 415 416 that the connection between local climate and large-scale ISM dynamics might be more complex. The YD is clearly defined as the cluster with heaviest $\delta^{18}O_{IVC}$ values, 417 418 suggesting a more proximal moisture source with little freshwater influence from 419 riverine runoff, and an overall weakened ISM circulation (Fig. 4B). Local hydroclimate conditions (indicated by δ^{13} C), on the other hand, appear to have been 420 421 rather similar to those during the preceding B-A, but more arid than during the 422 succeeding Holocene. This apparent inconsistency (weaker ISM, without increased 423 aridity at local level) reflects the different controls on $\delta^{18}O_{IVC}$ and $\delta^{13}C$, where 424 changes in the moisture source and composition do not necessarily always influence 425 local infiltration directly (Cheng et al., 2016b). While $\delta^{18}O_{IVC}$ is influenced primarily by the ISM during summer months, δ^{13} C is more sensitive to dry conditions, i.e., the 426 arid winter months. However, the dry season months in Meghalaya are characterized 427 428 by very dry conditions at present, and it is unlikely that conditions during the YD 429 were much different (as drier than dry is impossible). It is thus likely that a change in 430 precipitation seasonality during the YD led to a weaker ISM with a more proximal rainfall source during the summer months (i.e., heavier $\delta^{18}O_{IVC}$), and at the same time 431 432 a more even distribution of rainfall over the year, resulting in reduced seasonality and 433 little effective change in karst processes (i.e., lighter δ^{13} C).

434

435 Comparison to other records from Mawmluh Cave

436 To test if MAW-6 indeed reflects climate variations and not just local effects we compare the MAW-6 $\delta^{18}O_{IVC}$ record to the KM-A record (Berkelhammer et al., 2012) 437 438 and the MWS-1 record (Dutt et al., 2015) from the same cave (Fig. 6). We find good 439 visual replication between the three records on centennial time scale when 440 recalculating the age models for KM-A and MWS-1 using COPRA (Fig. 6). The 441 absolute difference in δ^{18} O values, especially pronounced between MAW-6 and 442 MWS-1, is likely related to varying degrees of isotopic fractionation at different drip 443 sites in the cave (similar to e.g., Stoll et al., 2015). For more quantitative information, 444 the three time series were interpolated to annual resolution and low-pass filtered in 445 order to only consider centennial time scale variations. Correlations were then 446 calculated by downsampling the data to 50-year resolution. With this approach, we 447 find high positive correlations between MAW-6 and KM-A during the period 6.9-9 448 kyr BP (r = 0.93), as well as between 9-12.4 kyr BP (r = 0.78, discarding the time 449 periods corresponding to hiatuses in MAW-6). Similarly, correlation between MAW-6 and MWS-1 is positive (r = 0.89). All relationships are highly significant ($p < 10^{-9}$). 450 Overall, this comparison corroborates our interpretation that variations in $\delta^{18}O_{IVC}$ in 451 452 MAW-6 are driven by climate. The high resolution and the precise chronology of our 453 record could significantly improve the available data from the ISM realm.

454

455 Comparison to other AM records

456 We chose three high-resolution and precisely dated records from Chinese caves 457 (Dongge, Dykoski et al., 2005; Yamen, Yang et al., 2010; Kulishu, Ma et al., 2012), 458 the MWS-1 record from Mawmluh Cave, and the NGRIP ice core record from 459 Greenland (Andersen et al., 2004) to compare with our MAW-6 $\delta^{18}O_{IVC}$ record (Fig. 7). Comparison of the MAW-6 $\delta^{18}O_{IVC}$ record to these reconstructions reveals very 460 461 similar centennial-millennial scale trends over the last deglaciation, further 462 corroborating our interpretation of the record as a proxy for ISM strength (Fig. 7). 463 However, more subtle differences are apparent as well. For example, whereas the 464 other AM reconstructions indicate the weakest summer monsoons during the last 465 glacial (until ~14.5 kyr BP), reflecting the pattern found in Greenland ice cores, 466 MAW-6 records weakest ISM conditions during the YD (Fig. 7). This is partly related to the adopted correction for ice volume and SST, which results in lighter $\delta^{18}O_{IVC}$ 467

468 during the last glacial, but the pattern is also apparent in the original δ^{18} O record (Fig. 469 3). This is possibly a reflection of changes in regional seasonality in NE India, with a 470 less vigorous ISM fed from proximal moisture sources, together with a wider spread 471 of precipitation over the entire year during the YD. Testing this hypothesis requires 472 seasonally resolved time series with highly robust chronologies.

473

474 In addition, differences appear when comparing the B-A interstadial in MAW-6 to 475 other records. Whereas the AM records considered here all show a relatively rapid 476 transition at the beginning and the end of the interval, with a plateau of lighter δ^{18} O 477 values during the B-A (attributed to increasing insolation; Ma et al., 2012), MAW-6 478 shows a pattern of rapid isotopic depletion at ~14.5 kyr BP followed by a gradual 479 increase towards YD values that is more similar to the transition recorded in 480 Greenland ice cores (Fig. 7). This might hint towards a close connection between NE 481 India and the North Atlantic realm, driven by the Westerlies. Evidence from paleoclimate records from Central Asia suggests that the AM and Westerly climates 482 483 are tightly connected over glacial-interglacial cycles (Cheng et al., 2016b). Mawmluh 484 Cave is located close to the Tibetan Plateau, with frequent influence of dry air masses 485 from the Tibetan High during the winter season, and a closer connection to the 486 Westerly climate than found at the Chinese cave sites is thus plausible. However, this 487 interpretation needs to be cautiously evaluated, due to limited replication with the 488 other δ^{18} O record from Mawmluh Cave covering the B-A interval (MWS-1; Dutt et 489 al., 2015), possibly related to chronological uncertainties in both records at this time.

490

491 The transition into the Holocene in MAW-6, although interrupted by a hiatus, shows 492 substantially lighter δ^{18} O values and thus ISM strengthening over time, similar to the record from Yamen Cave (Yang et al., 2010) and well-replicated in MWS-1. 493 494 Conversely to the gradually lighter δ^{18} O values found in the Yamen and Dongge cave 495 records, however, both reconstructions from Mawmluh Cave show a short plateau of 496 intermediate values between ~10.2 and 10.8 kyrs BP. In MAW-6, this interval is 497 demarcated by two hiatuses, and therefore direct comparison to other records is 498 difficult. This feature in the δ^{18} O records could reflect slow retreat or even a short-499 lived advance of the Himalayan glaciers, related to the increase in moisture and 500 precipitation (strengthening ISM) at the onset of the Holocene (Meyer et al., 2009).

501 The glaciation in the mountain range and related cold air outflow from the Himalaya 502 mountain range would have hampered the intrusion of the ISM somewhat longer in 503 NE India. This explanation remains hypothetical however, especially due to the 504 scarcity of data from the region.

505

506 **Dynamical changes in the ISM**

507 Dynamical regime changes in the ISM between late glacial and early Holocene were 508 investigated using recurrence quantification analysis (Ozken et al., 2015; Eroglu et al., 509 2016). We find a significant regime transition between LG and EH ISM, with more 510 chaotic conditions during the LG, but higher predictability during the EH (Fig. 8). 511 Disruption and weakening of the ISM during the LG, with frequent influence from 512 Westerly air masses and the Tibetan High, very likely result in less predictable 513 conditions. This is similar to findings from complex network analysis of the AM, 514 where weaker supra-regional links were found during the cold/dry Little Ice Age 515 (100-400 yr BP), suggesting that a weaker ISM is less predictable on a regional scale 516 (Rehfeld et al., 2012). During the EH, on the other hand, the strong seasonality 517 induced by the ISM would lead to more regular annual cycles in precipitation, and to 518 a higher predictability.

519

520 CONCLUSIONS

521 Stalagmite MAW-6 provides new paleoclimate data from Mawmluh Cave in NE India, covering the last deglaciation. We combine decadal-scale δ^{18} O and δ^{13} C 522 523 measurements on MAW-6 to unravel climate change at regional and local scale over 524 this period. A substantial post-glacial shift towards more negative δ^{18} O values is 525 interpreted as strengthening of the ISM, with maximum expression during the early 526 Holocene. This pattern is in agreement with other reconstructions from Mawmluh 527 Cave and the AM realm. Both the B-A and YD periods are clearly demarcated in the 528 record as stronger and weaker ISM, respectively. δ^{13} C is interpreted as reflecting local 529 hydroclimatic conditions, and is generally similar to δ^{18} O, suggesting that a 530 weak/strong ISM results in drier/wetter conditions at the study site. An intriguing 531 exception to this rule is the YD, where combined δ^{18} O and δ^{13} C analysis suggests a 532 reduction in precipitation seasonality, together with weakening of the ISM. Statistical 533 time series analysis of the δ^{18} O record reveals a significant regime transition over the

last deglaciation, with less predictable ISM during the late glacial, and higher
predictability during the Holocene, which we relate to the build-up of strong
precipitation seasonality induced by the ISM.

537

538 ACKNOWLEDGMENTS

539 We gratefully acknowledge financial support from the Swiss National Fond (SNF 540 Sinergia grant CRSI22 132646/1 and grant P2EZP2_172213), the German Science 541 Foundation (DFG project MA4759/8-1: Impacts of uncertainties in climate data 542 analysis (IUCliD): Approaches to working with measurements as a series of 543 probability distributions, and grant no. RE3994-1/1), the Chinese NSFC grants 544 4123054 and 2013CB955902, the US NSF grant 1103403, and the European Union's 545 Horizon 2020 Research and Innovation programme under the Marie Skłodowska-546 Curie grant agreement No 691037 (QUEST). We thank our Indian colleagues Bijay 547 Mipun and Gregory Diengdoh for their logistical help. We thank Daniel Gebauer for 548 support during fieldwork. We also thank Lydia Zehnder and Stewart Bishop (both at 549 ETH Zürich) for assistance during XRD and stable isotope analysis, respectively. Tim 550 Eglinton is acknowledged for financial support of F.A.L. We thank Ashish Sinha, 551 Max Berkelhammer, James Baldini, Yanjun Cai, and two anonymous reviewers for 552 constructive feedback and fruitful discussions on this and earlier versions of this 553 manuscript. We thank the editors Matthew Lachniet and Lewis Owen for feedback 554 and handling of the manuscript.

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805 Tables

806 Table 1: U-series dating results for stalagmite MAW-6. The errors given are 2σ . $^{*}\delta^{234}U = ([^{234}U/^{238}U]_{activity} - 1) \times 1000. ^{**}\delta^{234}U_{initial}$ was calculated based on ²³⁰Th age (T), 807 i.e., $\delta^{234}U_{initial} = \delta^{234}U_{measured} \times e^{\lambda 234xT}$. Corrected ²³⁰Th ages assume the initial ²³⁰Th/²³²Th 808 atomic ratio of $4.4\pm2.2 \times 10^{-6}$. Those are the values for a material at secular 809 810 equilibrium, with the bulk earth ²³²Th/²³⁸U value of 3.8. The errors are arbitrarily 811 assumed to be 50%. ***B.P. stands for "Before Present" where the "Present" is defined 812 as the year 1950 A.D. Values are indicated at one decimal place more than significant, 813 to avoid rounding errors. Ages excluded from the final chronology are shown in 814 italics.

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816 Figures

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Figure 1: A) Map with summer climatological conditions in the broader study area. 818 819 The location of Mawmluh Cave in NE India is indicated by the black dot. Other 820 discussed cave locations are indicated by the grey dots and arrows (D: Dongge Cave, 821 Y: Yamen Cave, K: Kulishu Cave). The dashed line indicates maximum northward 822 extent of the Intertropical Convergence Zone (ITCZ), which drives monsoonal 823 circulation. Brown arrows delineate dominant ISM circulation patterns, Asian 824 Summer Monsoon (ASM) winds are shown in green. B) Map of Mawmluh Cave. 825 Stalagmite MAW-6 was found broken in the West Stream (map courtesy of Daniel 826 Gebauer). C) Scan of cut and polished stalagmite MAW-6.

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Figure 2: Age-model of MAW-6 (constructed using cubic interpolation in COPRA).

830 The median of the age model is shown in blue, with the 95% confidence intervals in

831 light grey. The U-series ages used to construct the age model are shown in black,832 while the excluded ages are in red. Hiatuses are indicated by dashed black lines.

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Figure 3: A) δ^{18} O and δ^{13} C records with median and 95% confidence intervals. The time periods discussed in this study are indicated at the bottom of the figure (IM: Intermediate period, 10.8-10.2 kyr BP, YD: Younger Dryas, B-A: Bølling-Allerød, LG: late glacial). Major controls on δ^{18} O (ISM strength) and δ^{13} C (amount of in-cave fractionation due to cave air pCO₂ and drip rate) are indicated by the bars.

B) Cross-correlation between δ^{18} O and δ^{18} O_{IVC} vs. δ^{13} C, estimated using kernel-based cross correlation analysis (Rehfeld and Kurths, 2014) with the toolbox NESTool (<u>http://tocsy.pik-potsdam.de/nest.php</u>).

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Figure 4: A) δ^{13} C vs. ice volume corrected δ^{18} O (δ^{18} O_{IVC}) relationship in stalagmite 845 846 MAW-6. The record can be subdivided into clusters corresponding to different time 847 periods: Holocene (Hol), Intermediate (IM), YD, and B-A and late glacial (B-A/LG). All clusters show high degrees of correlation between $\delta^{13}C$ and $\delta^{18}O_{IVC}$, indicated by 848 849 the corresponding r values (same values as in Fig. 3B). Arrows indicate the direction of the main forcings (dry season aridity and ISM strength). B) Boxplots for $\delta^{18}O_{IVC}$ 850 851 and δ^{13} C (top and bottom panel respectively). Boxes are defined by the median (red 852 line) and delimited by the 1st and 3rd quartile. Whiskers define the lowest and highest 853 value within 1.5 times the inter quartile range of the cluster. Outliers are indicated by 854 red crosses.

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Figure 5: Schematic of the factors influencing isotope signals at Mawmluh Cave. Data is derived from monitoring studies at the cave site (Breitenbach et al., 2015), and precipitation data is from the Indian Meteorological Department Station Cherrapunji.

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862 Figure 6: Comparison of stalagmite δ^{18} O records MAW-6 (blue), MWS-1 (orange,

B63 Dutt et al., 2015) and KM-A (purple, Berkelhammer et al., 2012). Proxy uncertainties

864 (95% confidence intervals), as calculated by COPRA, are shown in light shading.

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Figure 7: Comparison of MAW-6 $\delta^{18}O_{IVC}$ to $\delta^{18}O$ records from the ISM and broader AM regions, as well as to the NGRIP Greenland ice core record. MAW-6 reflects and corroborates previous reconstructions from the Asian monsoon region showing the weakest ISM after the deglaciation occurring during the Younger Dryas, and stronger ISM during the preceding Bølling-Allerød, as well as during the Holocene.

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Figure 8: Results of the TACTS analysis on MAW-6 δ^{18} O. The cumulative probability distribution established through 5000 random realizations of DET measure is shown by the blue line, with gray shading indicating the 95% confidence interval. The late glacial (LG) and early Holocene (EH) are characterized by distinct DET measures (0.663 and 0.736, respectively). While the LG is within the 95% confidence interval, the EH is outside, highlighting the high predictability of the ISM during this period.

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